

# Technologies for a future giant optical telescope

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*Opportunities for international research cooperation and commercialisation of  
functional materials  
Helsinki, 22 January 2008*

# Technologies for a future giant optical telescope

## ❖ Talk summary

- What is ESO?
  - Background to the European Extremely Large Telescope (E-ELT)
  - Present activities
  - Time scales
  - Technology examples
- ❖ Work of ESO and many collaborating institutions and industrial partners



## **ESO is ...**

- **Inter-governmental Organisation**
- **Established by International Convention in 1962**

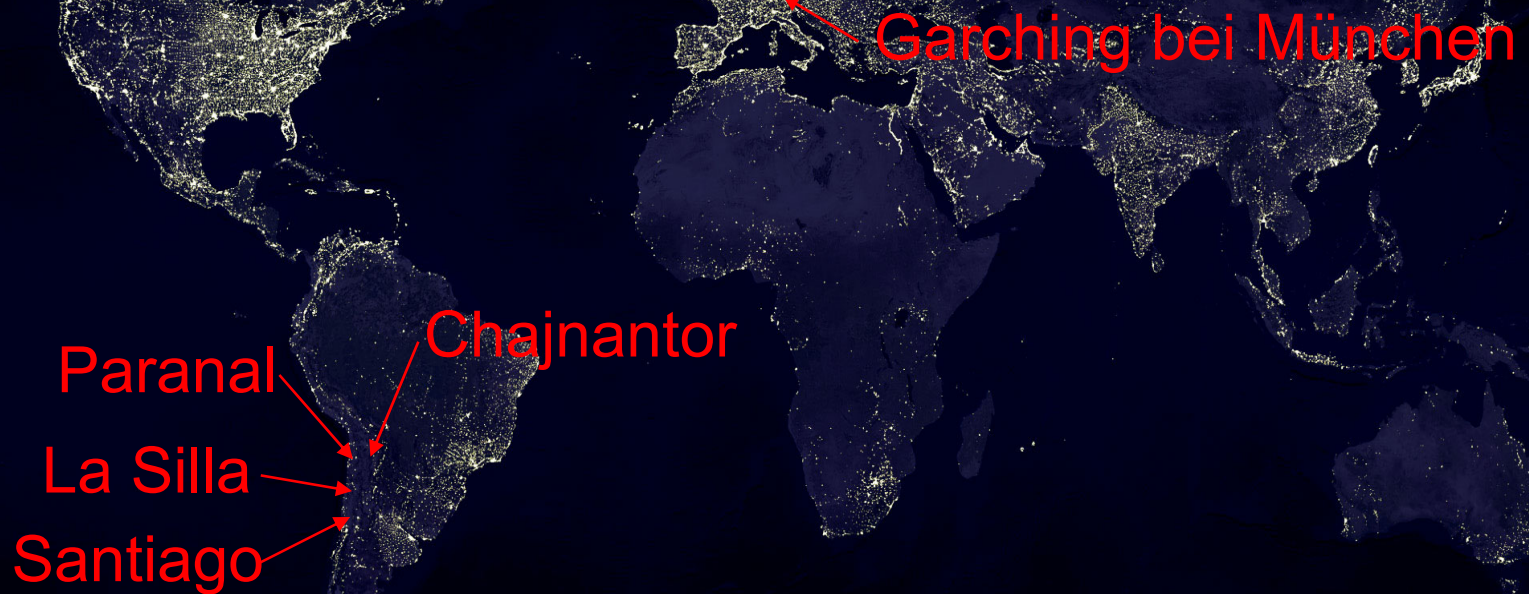
## **ESO's mission...**

- **Provide astronomers in member states (~ 3000) with state-of-the-art observational facilities**
- **Further and organise collaboration with astronomy and astrophysics in Europe**

## **Budget and Staff ...**

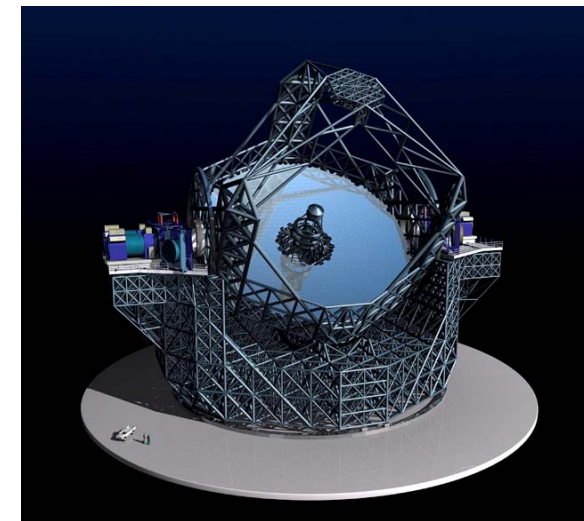
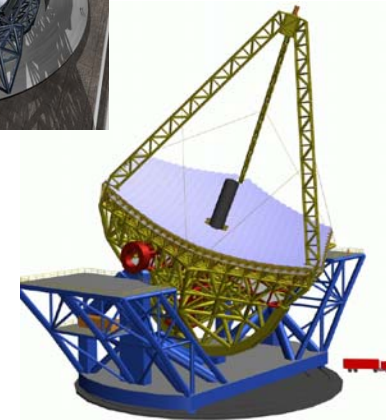
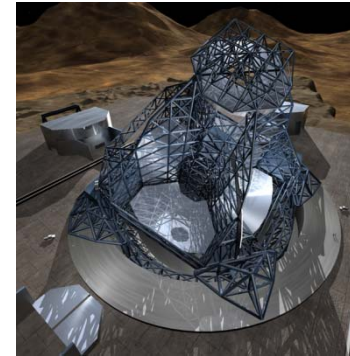
- **Budget: ~200 Euro**
- **Staff: 650 (2006)**

# ESO Sites



# E-ELT Background

- ❖ OWL 100m Design Study (1997 - 2005)
  - Why have telescopes only increased by a factor of two in diameter every 50 years?
  - What are the physical limits?
  - Can we break the cost scaling laws ( $D^{2.6}$ ) ?
  - Can a 100m optical telescope be built for 1000 M€ ?
- ❖ Swedish 50m telescope design study
- ❖ OWL Concept review (November 2005)
  - No show stoppers but high risk of not being completed on time on in budget
  - Recommended to explore the possibility of building a 30 – 60 m telescope at a cost of 650-750 M€ in a competitive time scale
  - 5 Working groups set up (Science, Instruments, Site, Design and Adaptive Optics)
- ❖ E-ELT baseline design accepted at the Marseille meeting (November 2006)
  - 42m diameter, 5-mirror design
  - Approval to start Phase B



# Present activities (1)

## FP6 Activities

- Technology oriented, design independent
- 30 partners, 11 countries, 30.5 M€ (8.4 M€ from FP6)
- ❖ **ELT Design Study (EC Research Infrastructures Action)**
  - Mechanical studies
  - Enclosure concepts and wind studies
  - Adaptive optics
  - Integrated modelling
  - Wavefront control
  - Advanced adaptive optics
  - Optical fabrication
  - Instrumentation (point designs)
  - Site characterisation
- ❖ **Other FP6 activities (JRA )**
  - Second generation Adaptive Optics
  - Fast Optical Detectors and Fast readout high performance optical detectors
  - Smart focal planes and VPG development

# Present activities (2)

## ESO Funded Activities (ESO and collaborating institutions)

### ❖ E-ELT Phase B

- Requirements definition, system engineering
- Mechanical design (internal and industrial studies)
- Industrial study contracts
- Optical design, adaptive relay unit industrial studies (M4 & M5)
- optical manufacture studies

### ❖ VLT Activities

- VLT Adaptive Optics Facility (inc. DSM and 4-laser guide star system)
- Second generation VLT Instruments, inc. SPHERE (exo-planet finder with extreme AO), MUSE/GALACSI & HAWK-I/GRAAL (with ground layer AO)

### ❖ Other In-house developments and studies

- Electronics standards, time standards, product testing (incl. cryo testing), control architectures, technology demonstrators, ...

### ❖ Other industrial and academic studies and developments

- Fibre lasers, Laser components, PC hollow core fibres, ...

# E-ELT Schedule

## ❖ Time scales

### ➤ E-ELT Phase-B

- Fully funded (~ 50 M€)
- Jan 2006 – December 2010
- In 2010:
  - All FP6/7 studies completed
  - Design concept agreed
  - Site agreed
  - All important technical specifications defined

### ➤ Construction phase: 2011 – 2018

- Subject to funding being available
- First science 2015 - 2016
- Timescale compatible with JWST

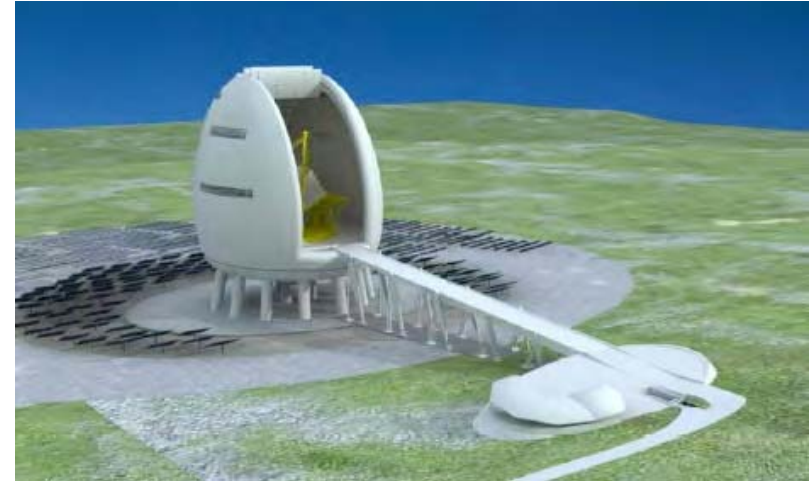
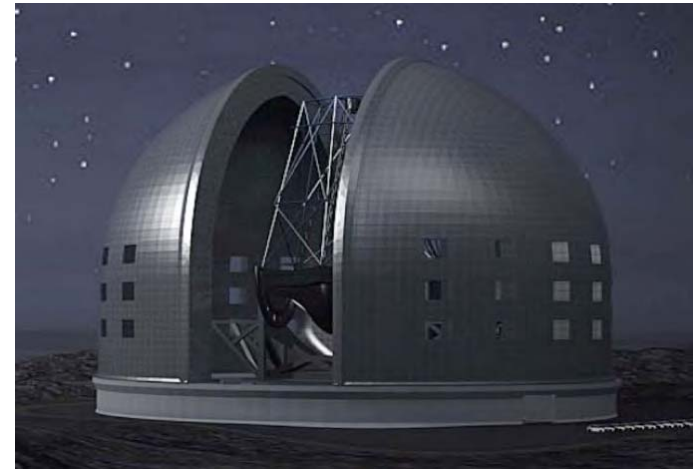


# Some key developments & technologies

- ❖ Enclosure concept studies
- ❖ Mirror Actuators
- ❖ Mirror Metrology
- ❖ Wind analysis & control
- ❖ Mirror Phasing (APE)
- ❖ Friction and other drive systems
- ❖ Optical fabrication
- ❖ Adaptive optics
- ❖ Fast optical detectors
- ❖ Fibre lasers
- ❖ Instrumentation developments & concepts

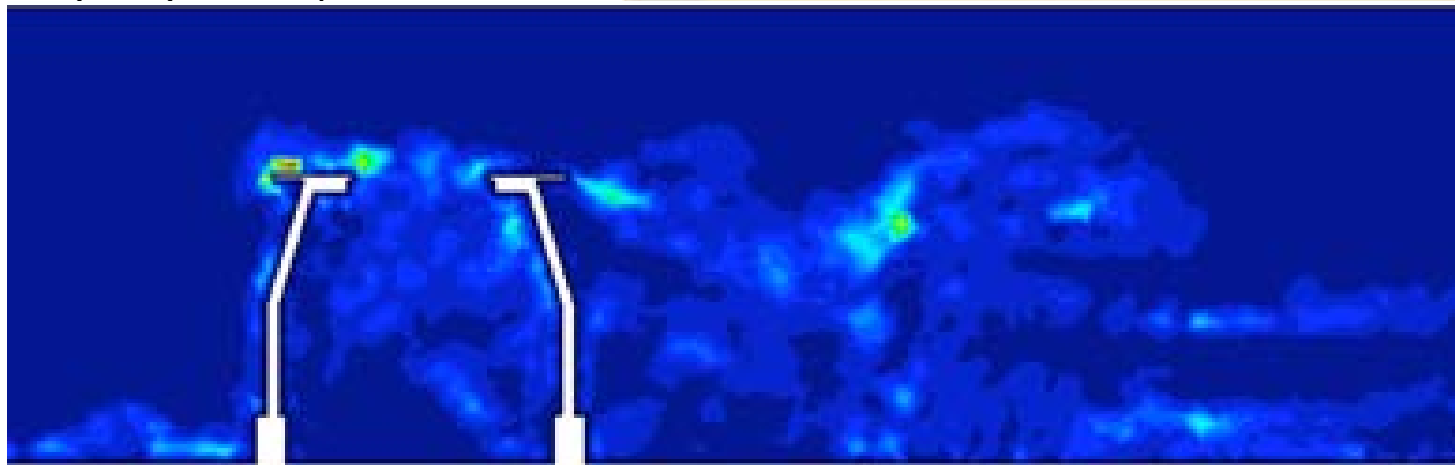
# Enclosure concept studies

- ❖ 4 concepts evaluated  
(Sliding enclosure, eye-lid, e-BAT and box-type)
- ❖ Evaluated for cost, reliability, maintainability, protection from wind and stray light, safety, etc.



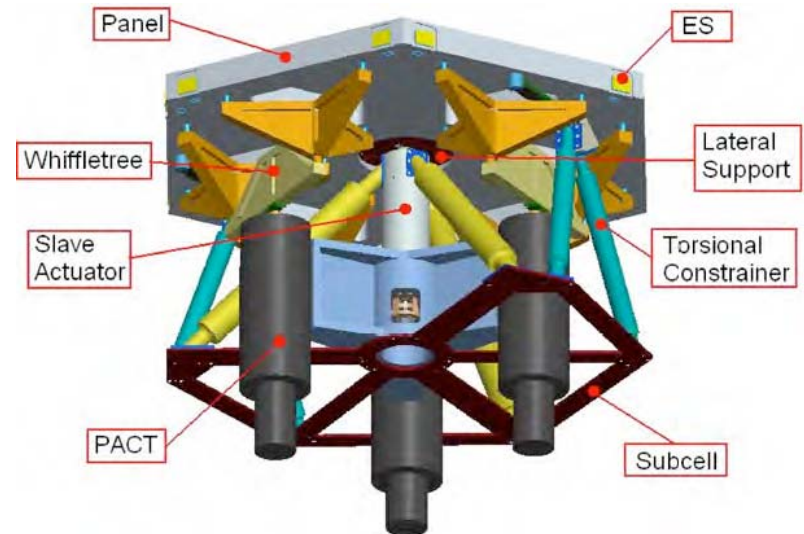
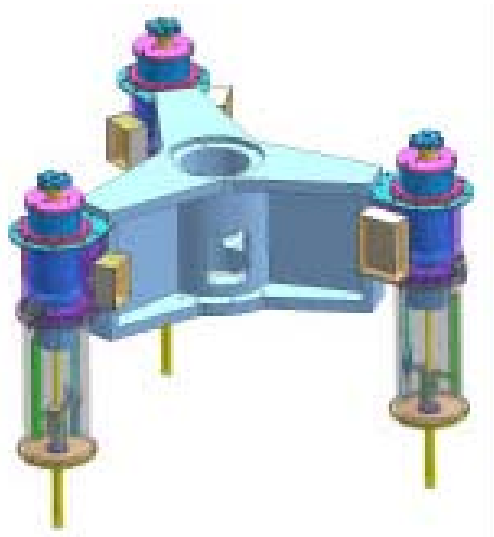
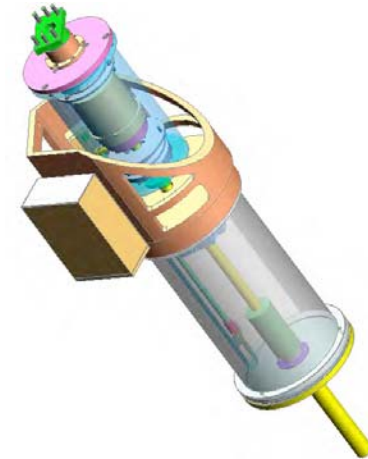
# Enclosure concept studies

- ❖ 4 concepts evaluated (Sliding enclosure, eye-lid, e-BAT and box-type)
- ❖ Evaluated for cost, reliability, maintainability, protection from wind and stray light, safety, etc.
- ❖ Turbulence evaluated with Large eddies simulator, CFD modelling and wind tunnel (ITER and Boundary Layer WT)



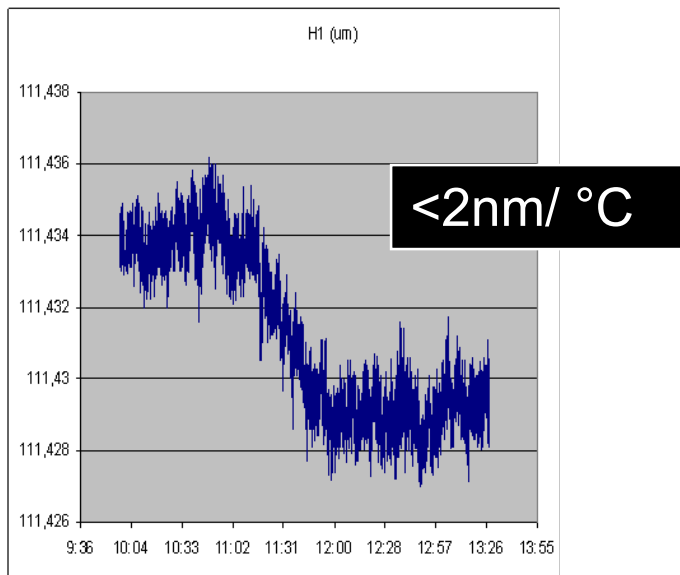
# Mirror Actuators

- ❖ E-ELT M1 position actuators designed and prototyped
- ❖ Dual-stage (voice-coil and stepper)
- ❖ Max stroke of 150mm for mirror segment removal
- ❖ Closed loop position accuracy of 5nm with variable load up to 1700 N
- ❖ Production of 18 units for WEB



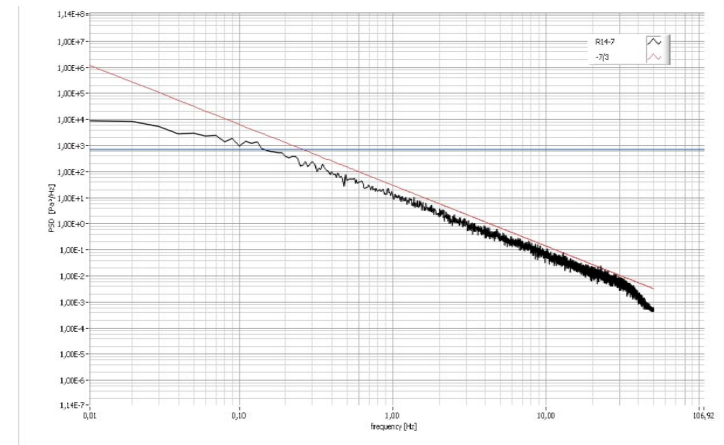
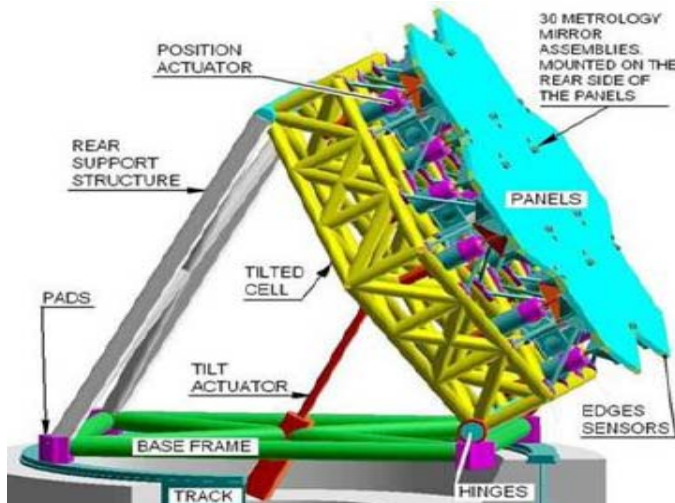
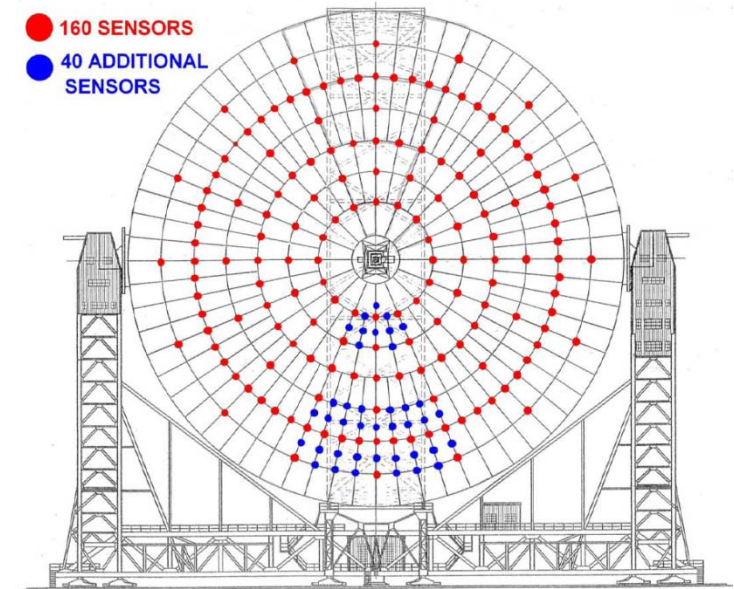
# Mirror metrology

- ❖ M1 mirror metrology
  - Capacitive and inductive sensors compared
  - Evaluation of stability, sensitivity to mirror misalignment and deformation of segments, cross talk, noise etc.)
  - High stability test bench has been developed
  - Hardware units produced for integration in the WEB project



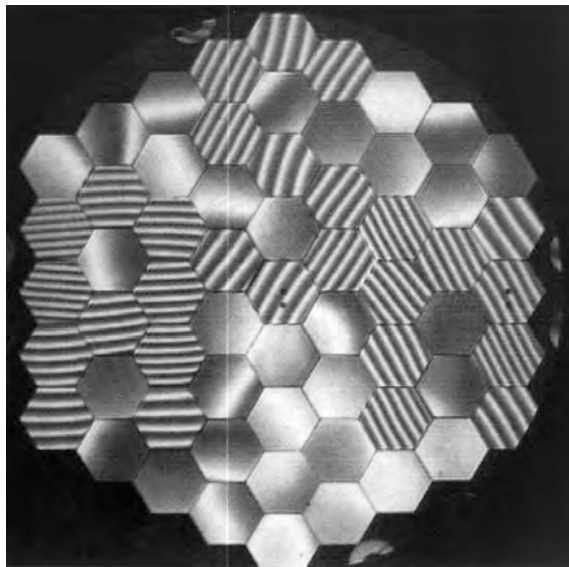
# Wind analysis and control

- ❖ **Jodrell Bank wind measurements**
  - 200 wind pressure sensors mounted on surface
  - 40 mounted closer together to measure short range correlation functions
- ❖ **VST enclosure measurements**
- ❖ **WEB (Wind Evaluation Breadboard)**
  - Dev. of position control architectures for wind rejection



# Active Phasing Experiment (APE)

- ❖ Designed to test 4 different phasing technologies for segmented mirrors (DISPSI, ZEUS, PYPS & SHAPS)
- ❖ A 61-element prototype constructed with internal metrology
- ❖ To be installed and tested on the VLT in 2008

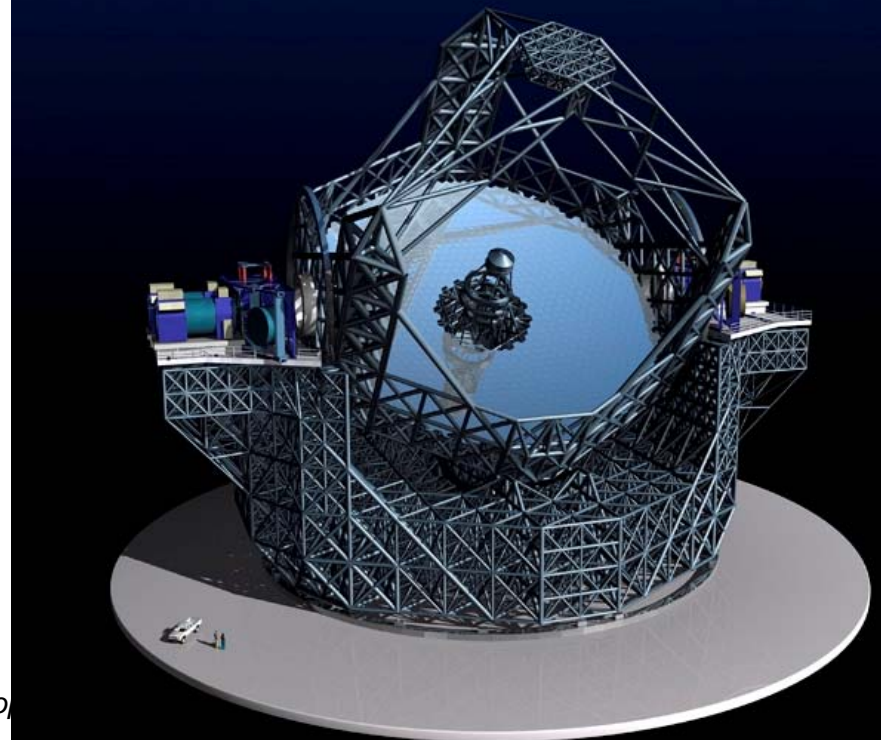


# Mechanical Design Studies

- ❖ OWL 100m concept design
- ❖ E-ELT 42m mechanical design
  - ESO Baseline design
  - Two industrial studies
- ❖ Mechanical technology studies (FP6)
  - Composite materials
  - Structural ropes
  - Magnetic drives
  - Magnetic levitation
  - Friction Drives >>



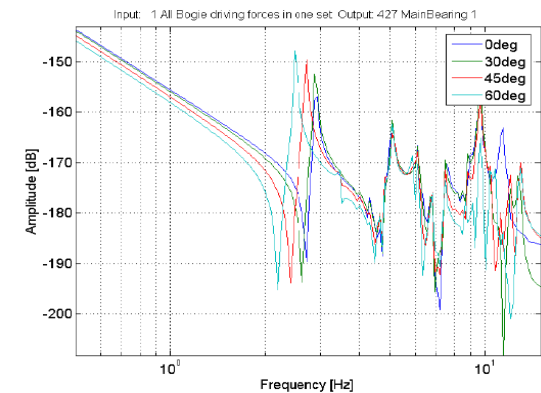
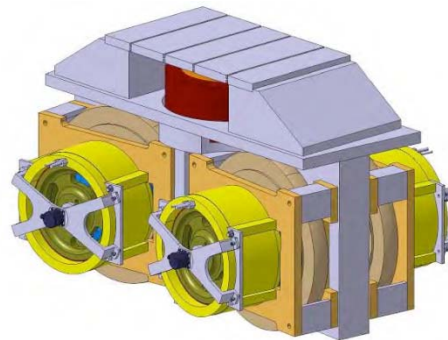
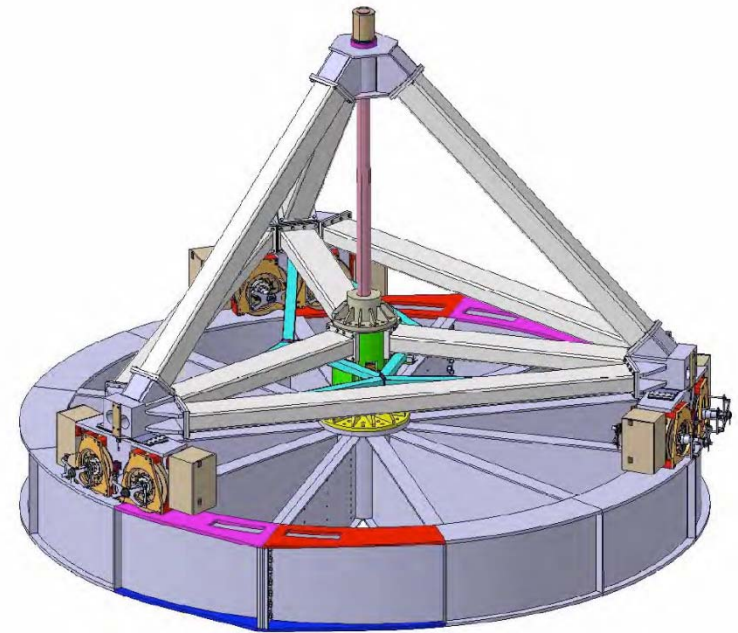
42m E-ELT





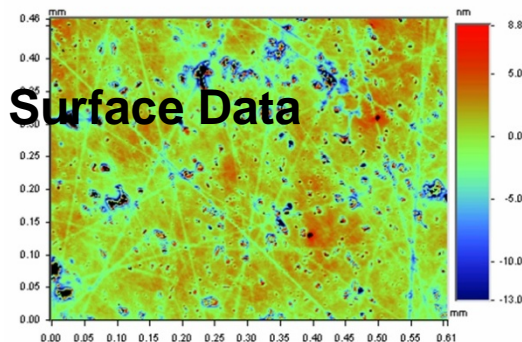
# Friction Characterisation Breadboard

- ❖ Concept evaluation for friction driven bogies
- ❖ 7m Diameter test rig, including wind disturbance and track misalignment generators
- ❖ Evaluation of stiffness, stick slip, tracking, pointing, sensitivity to track misalignments, sensitivity to wind disturbances and RAMS aspects
- ❖ Prototype not yet as stiff as required



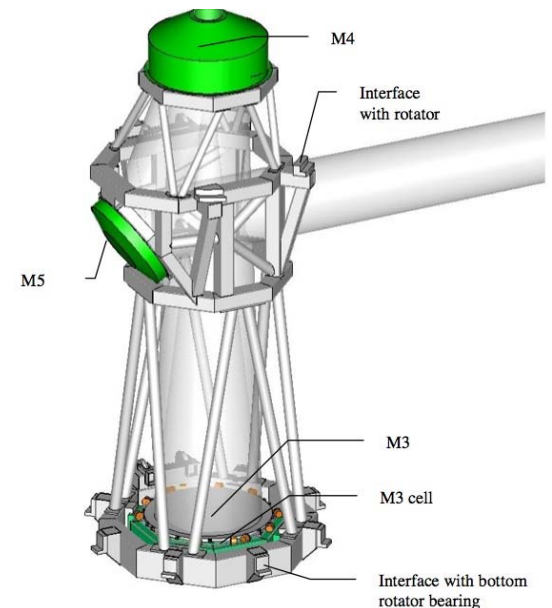
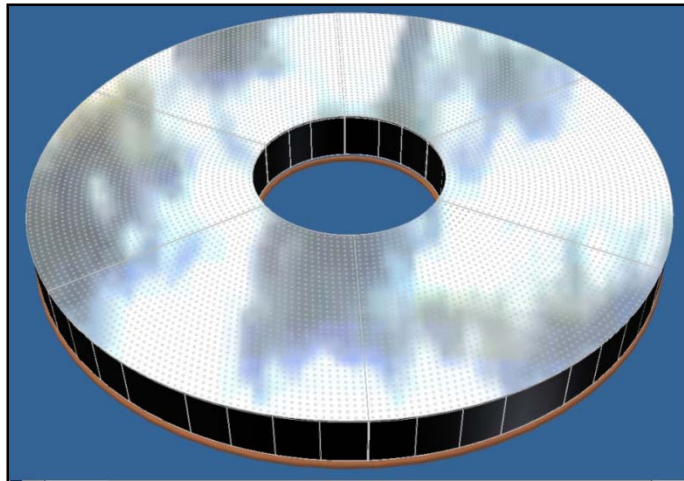
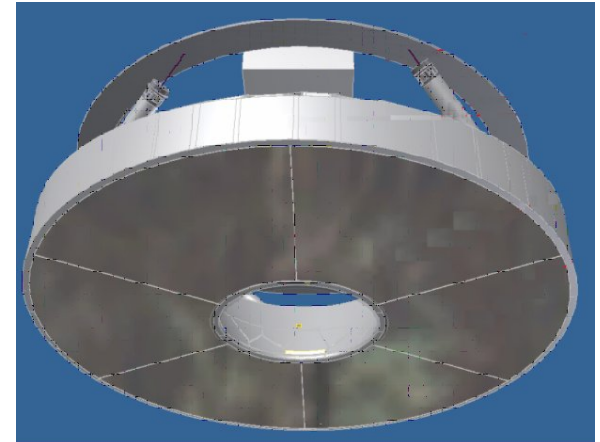
# Optical Fabrication Studies – M1

- ❖ Manufacture and testing of 1m-class SiC flat segments
- ❖ Qualification of SiC as an affordable alternative to Glass
- ❖ Investigation of over-coating technologies (CVI, ICVI, metallic, Si, ...)
- ❖ Investigation of polishing techniques (edge control)
- ❖ Production of flat segments for WEB
- ❖ ESO contract for the production of 7 figured segments in Zerodur, Astrosital, ULE + another material

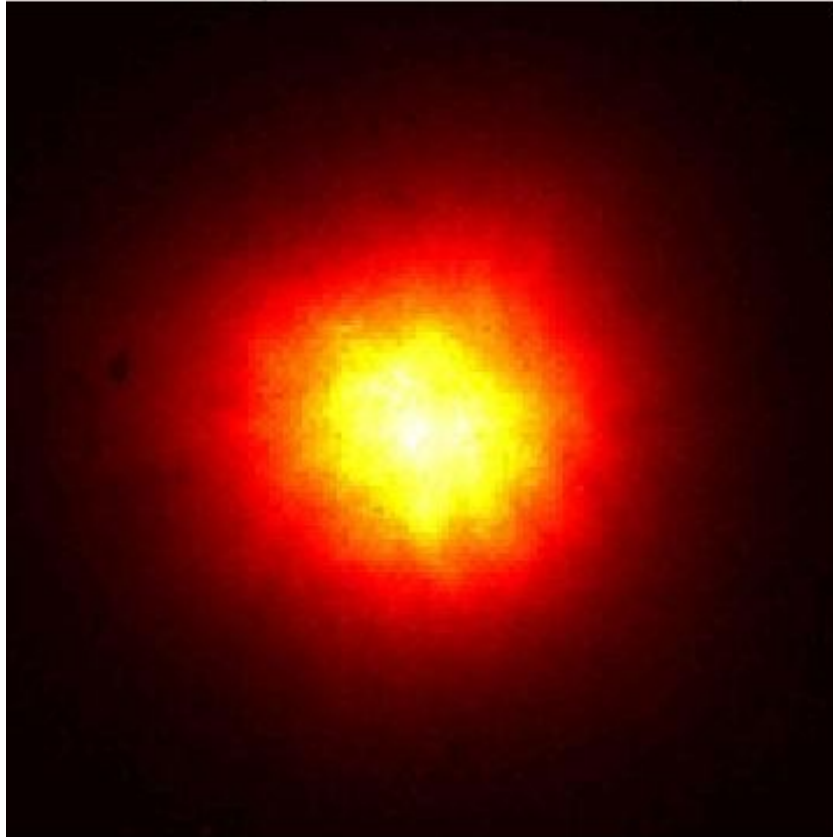


# Optical Fabrication studies M2 – M5

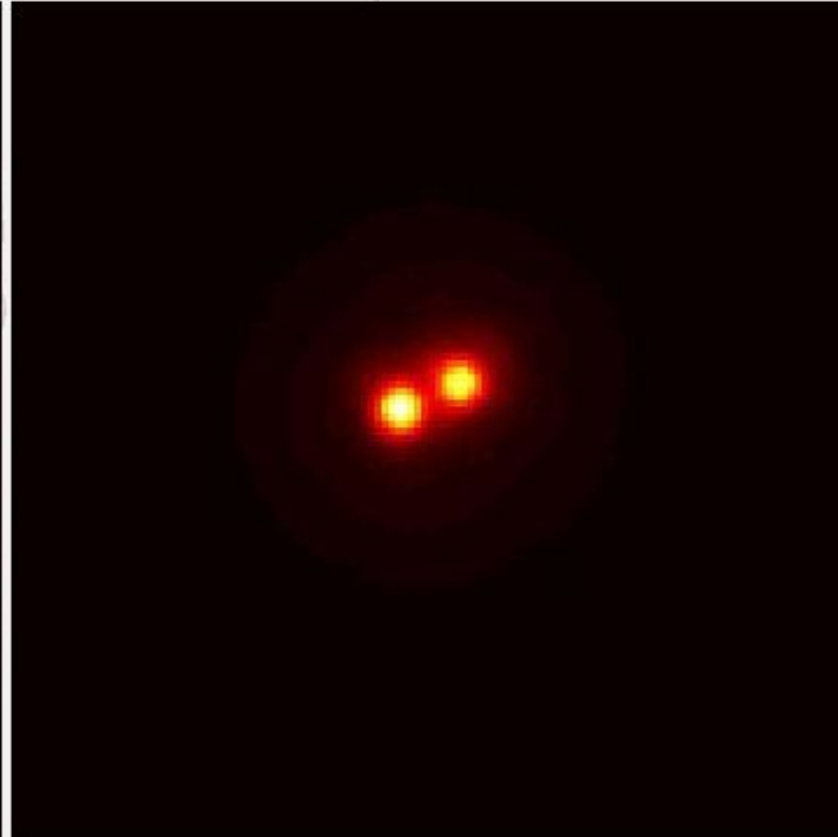
- ❖ Several industrial feasibility studies completed
- ❖ M2 6m convex (current technology)
- ❖ M3-M5 Adaptive Optics Relay Unit
  - M3 4.2m
  - M4 2.5m deformable
  - M5 2.7m tip-tilt
- ❖ VLT: 1.2 deformable M2 in manufacture



# Adaptive Optics



Non-AO



AO

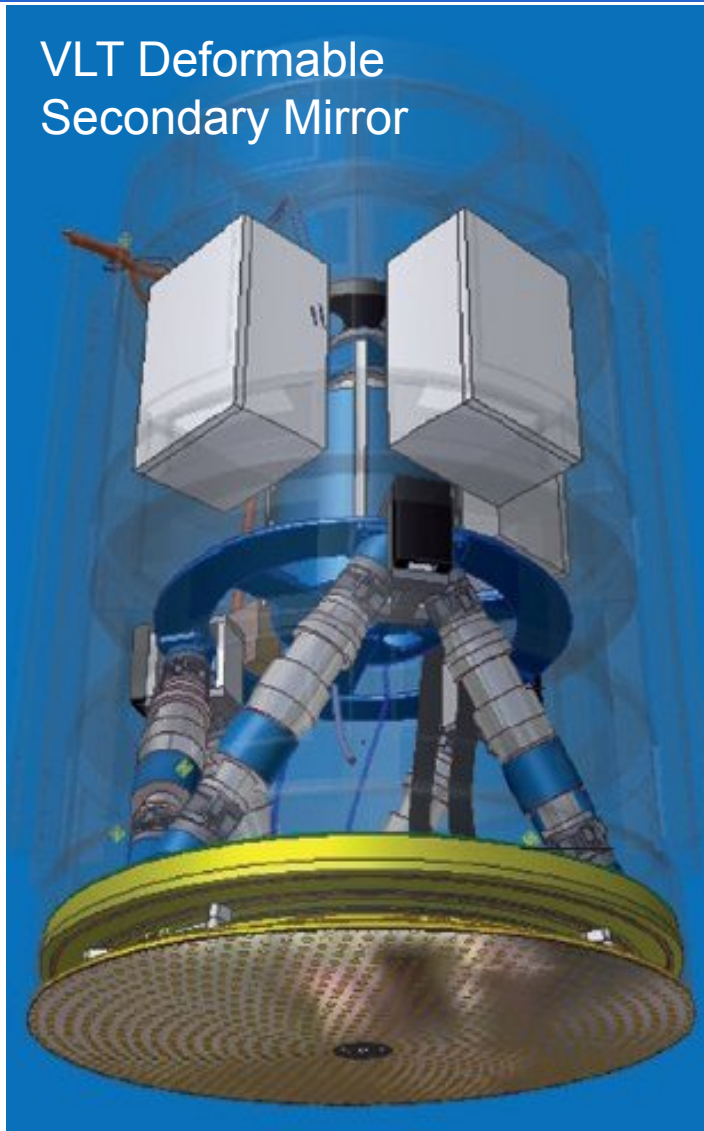
Picture from SINFONI on VLT: 0.75 arcsec seeing,  
binary separation 0.12 arcsec

*Technologies for a future giant optical telescope, Helsinki, 22 January 2008*

# Adaptive Optics

- ❖ **E-ELT AO Concept developments**
  - Single Layer AO (SLAO)
  - Ground Layer AO (GLAO)
  - Laser Tomography AO (LTAO)
  - Multi- Conjugate AO (MCAO)
  - Extreme AO (XAO)
- ❖ **Novel AO Concepts**
  - Dynamic refocus of a pulsed LGS (comparison of techniques, scaling to an ELT)
  - Centroid algorithms and static optical compensation (LGS Elongation)
  - Differential focal and spot anisoplanatism (Cone Effect)
  - Pseudo Infinite Guide Stars (PIGS)
- ❖ **Large, high density deformable mirrors**
  - 1371 Actuator DM for SPHERE
  - Deformable large thin shells
  - 1.2m Deformable mirror upgrade for M2 of VLT UT4 (Part of the VLT AO Facility) 1170 actuators, 29mm pitch, 80 Hz

VLT Deformable Secondary Mirror



# Overview of AO possible performance



	Category #1	Category #2	Category #3
<b>Corrected FoV (diameter)</b>	Small (Isopl. Angle, i.e. 20" with 0.8 of max Strehl)	Medium (1')	Large (5')
<b>Performance metric</b>	70% Strehl (K) 33% Strehl (J)	50% Strehl (K) 11% Strehl (J)	16%(K), 5% (J) EE in 50mas (4 x seeing)
<b>Obs. <math>\lambda</math> range</b>	R–N	J–N	J–N
<b>PSF uniformity</b>	low	high	medium
<b>Potential systems</b>	SCAO (1DM, 75% Strehl*, NGS, low sky coverage)	MCAO (2 DMs, several LGS)	GLAO (1 DM, several LGS, special LTAO mode)
	LTAO (1DM, 60% Strehl, several LGSs high 'sky coverage')	MCAO (3 DMs, several LGS), larger FoV	MOAO (1-2 DM, several LGS, , restricted to 'islands', better Ensquared Energy)
	XAO (2DM, bright NGS)		

Sky coverage always 100% (=> **LGS required**), not fulfilled by SCAO

Wavelength range R–N (deviations above due to performance)

All numbers in 0.5" seeing and with 42-m telescope

# Laser Guide Stars

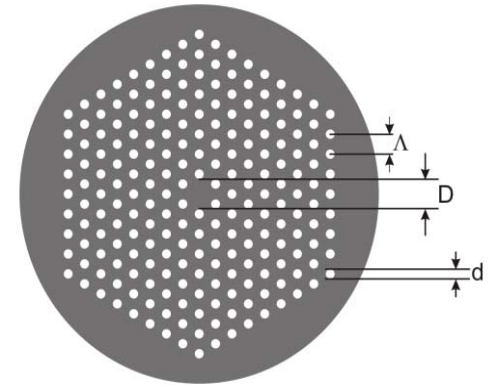
- ❖ **Current VLT LFSF**
  - 10 Watt CW 589 nm
  - 1.1 arcsecond artificial star
  - Fibre optical feed
  - In regular operation since spring 2007
- ❖ **Raman Fibre Lasers**
  - 1178nm frequency doubled
  - 5W currently achieved (7W specified), combined to achieve 20W power (15W launched)
  - Will be used in the VLT 4 LGSF
- ❖ **E-ELT pulsed lasers**
  - Planned for E-ELT
  - Needed to counter large ELT cone effect
  - On-going but delayed due to definition of system requirements



# Optical Fibres

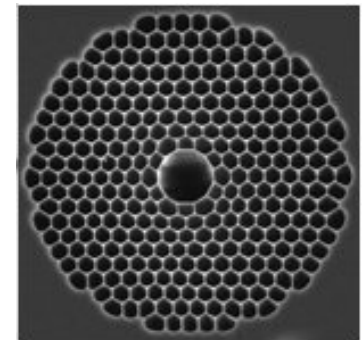
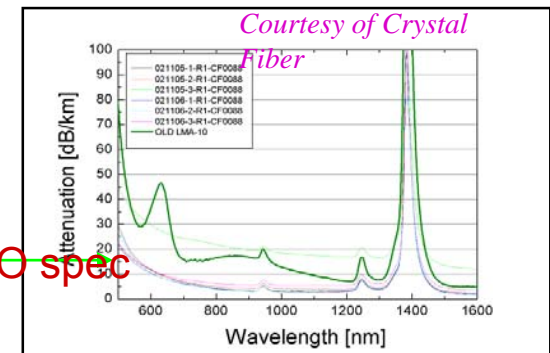
## ❖ Photonic Crystal fibres

- Current LGSF uses photonic crystal fibres produced by Crystal Fiber (DK) and Mitsubishi (J)
- Development contract to give acceptable transmission at 598nm



## ❖ Hollow core PC Fibres

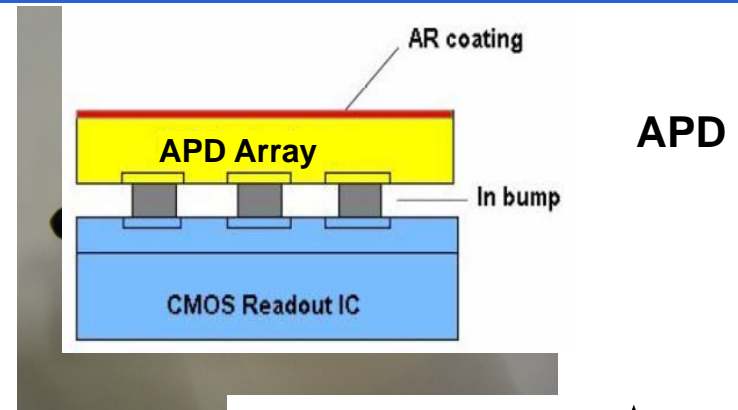
- Potentially allow greater laser power transmission
- Needed for pulsed lasers
- Development contract not successful due to pulling problems.



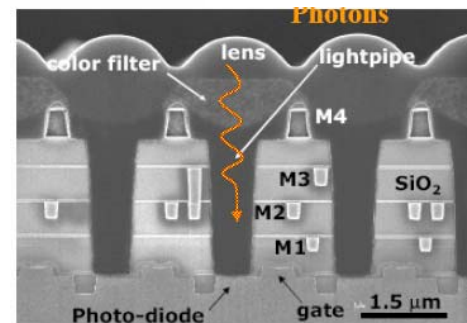
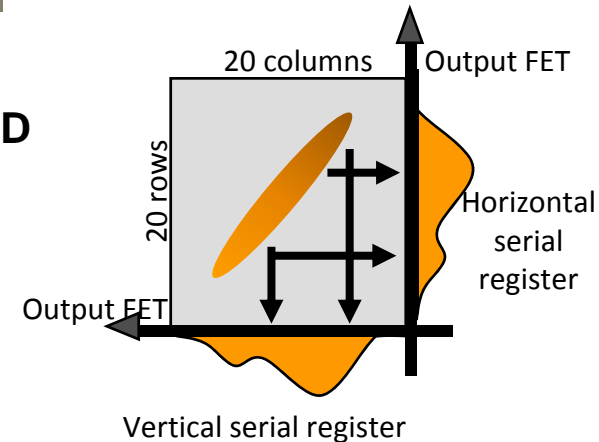


# Fast Detectors

- ❖ FP6 fast WFS project
  - 240x240 pixels
  - Based on E2V CCD220
  - Incorporated in OCAM test camera on GTC
  - To be used at ESO for SPHERE planet finder (2nd generation VLT instrument)
  
- ❖ Advanced HD WFS Detector
  - 4 concept studies carried out for a 1600x1600 pixel detector
  - RON  $< 3e^-$ , 700Hz frame rate
  - 3 different approaches proposed
  - Study completed end 2007
  - 2 technology validation studies to start 2008

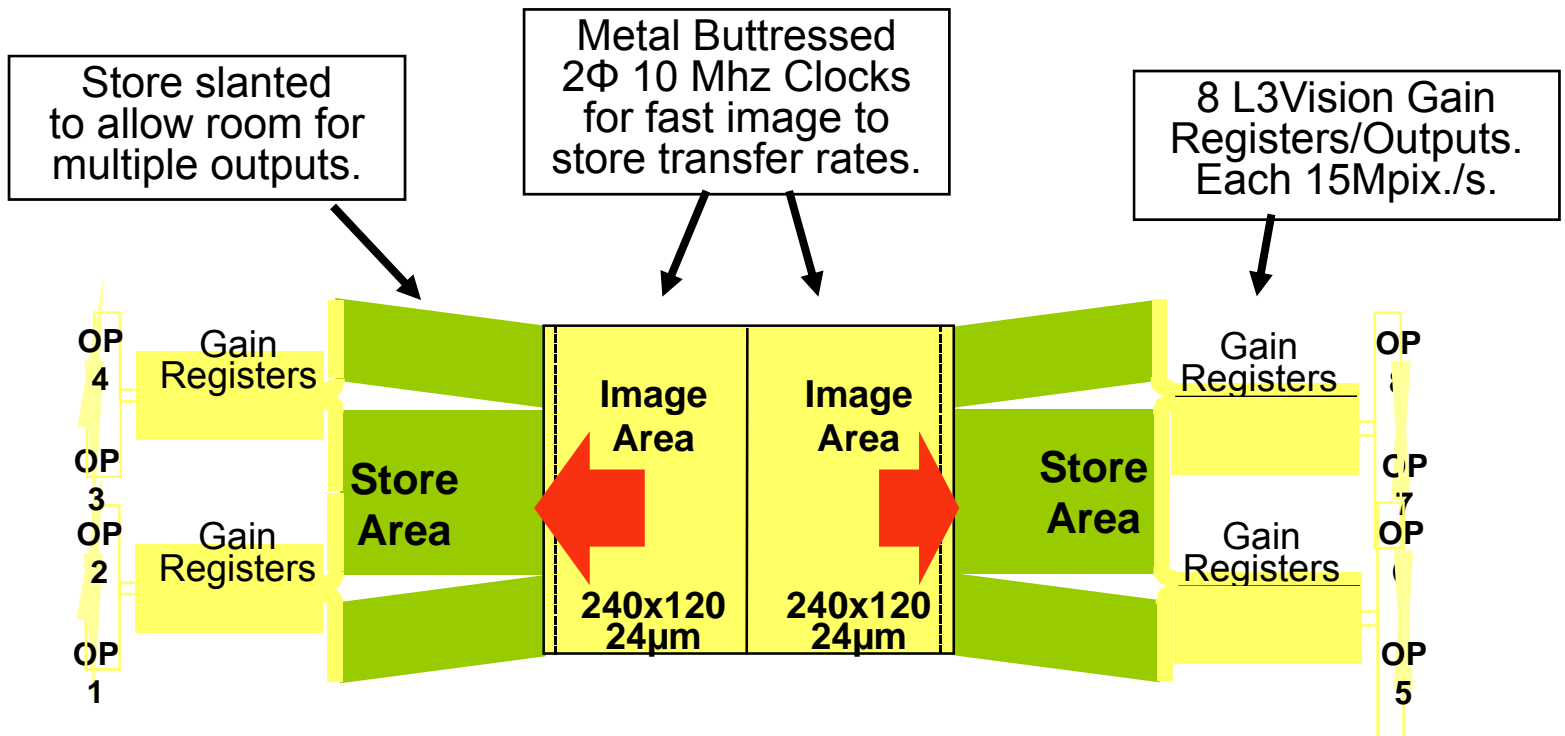


**Orthogonal transfer CCD**



**Monolithic CMOS**

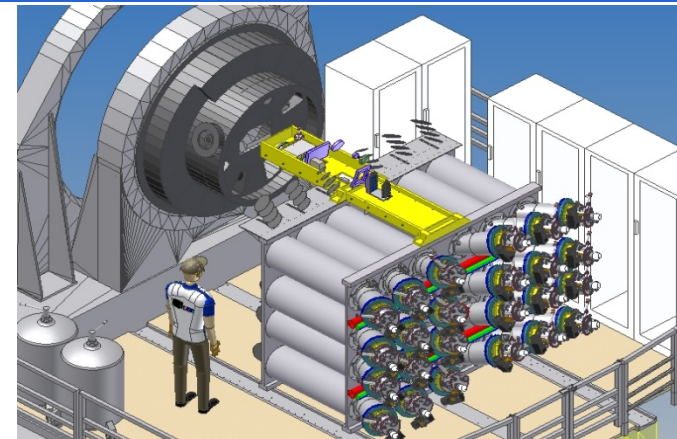
# The CCD 220 Design



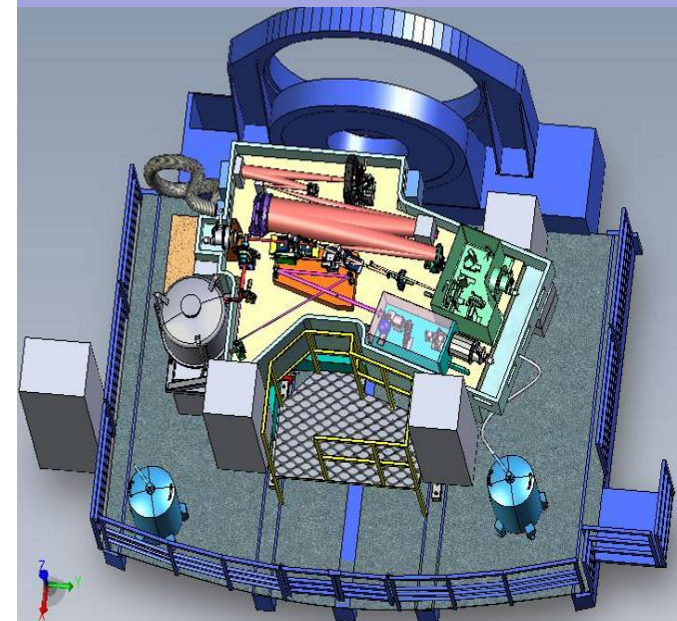
Split frame transfer 8-output back-illuminated e2v-L3Vision CCD.

# Instrumentation

- ❖ ESO/FP6 AO technology prototypes for incorporation into VLT instruments
  - MAD (Multi (dual) conjugate AO Demonstrator (In operation since mid-2007))
  - MUSE (GALACSI)
  - HAWK-I (GRAAL)
  - SPHERE (Planet finder instrument with eXtreme AO)
- ❖ FP6 concept studies for E-ELT instruments



**GALACSI-MUSE**

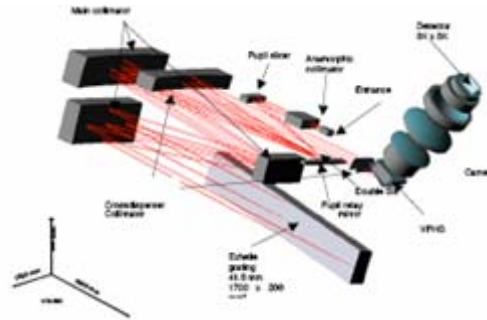


**SPHERE**

# Examples of E-ELT instrument concept studies



SCeLT



CODEX (HiSPEC Visual)



HiSPEC IR

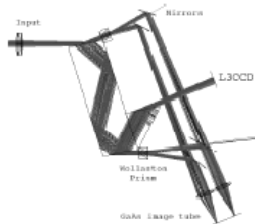
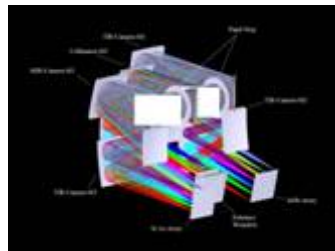
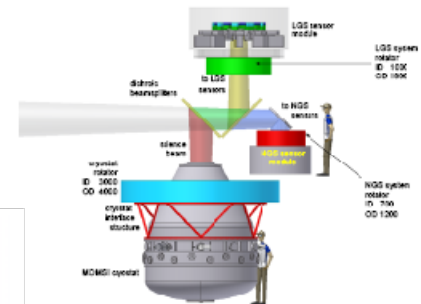


Figure B.2 CASP polarimeter, Zemax ray trace [6], courtesy of P. Collins

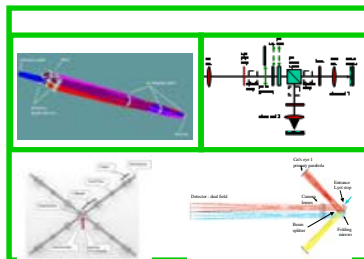
HiTri



Planet Finder - EPICS



MOMSI



MIDIR

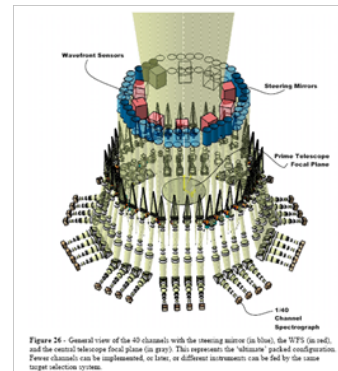
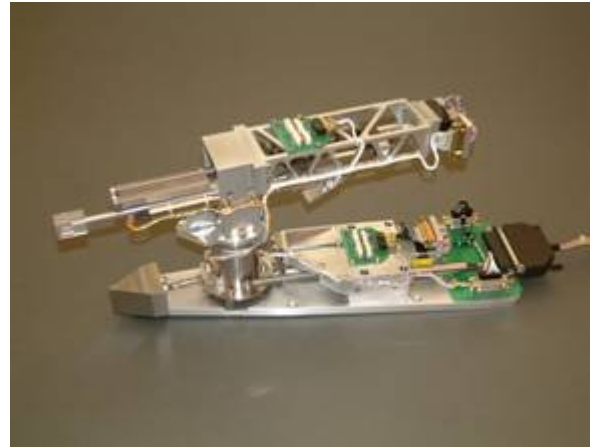
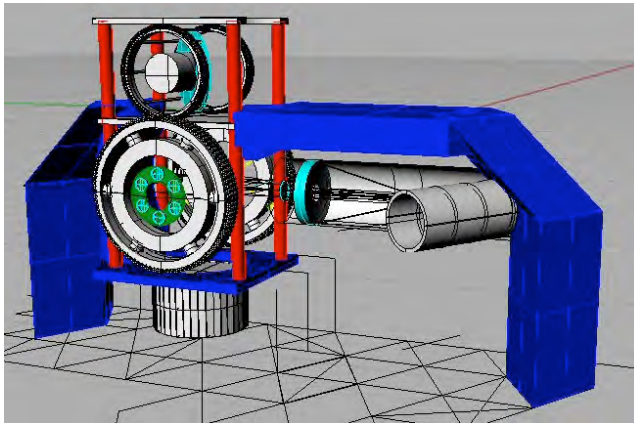


Figure 24 - General view of the 40 channels, with the mosaic matrix (in blue), the VFTs (in red), and the mosaic telescope focal plane (in grey). This represents the 'reference' packed configuration. Fewer channels can be implemented, or larger, or different instruments can be fed by the same target selection system.

# Many other activities ...



Smart focal planes



Subsystem design



Site testing