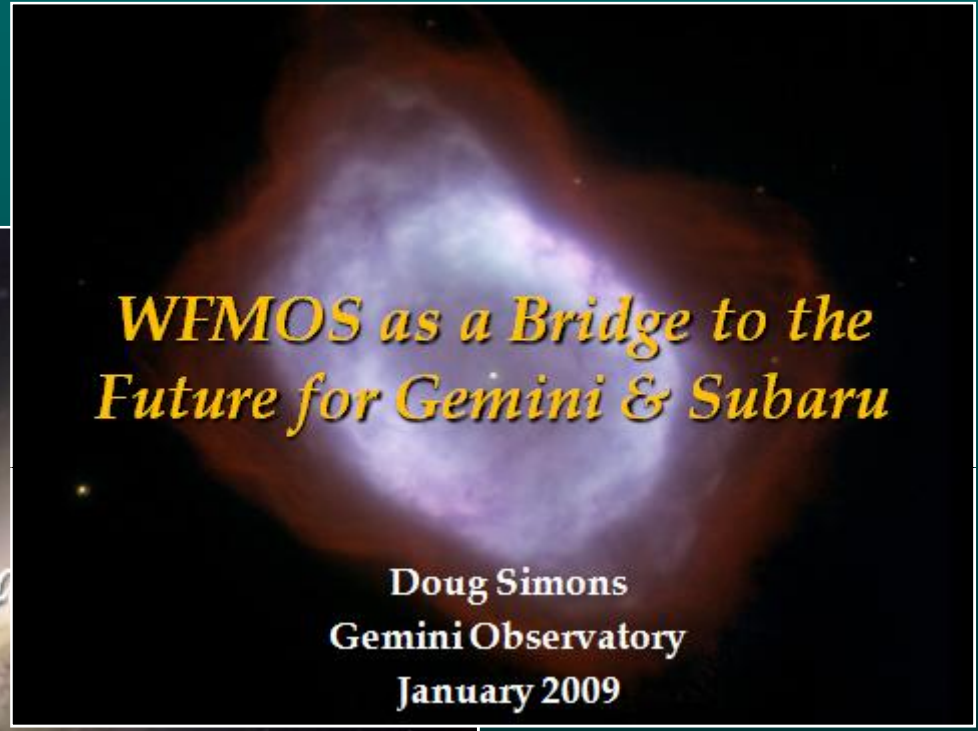




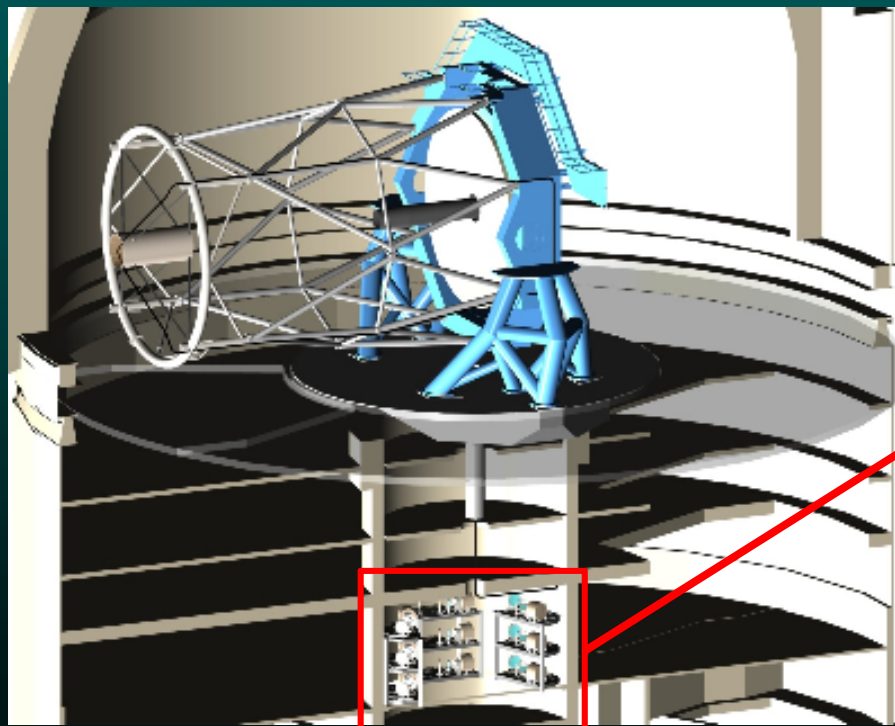
*WFEMOS as a Pathfinder into  
the Future of Astronomy*

Doug Simons  
Director, Gemini Observatory



## K.A.O.S.

Kilo-Aperture Optical Spectrograph





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# Onward to Aspen (June 2003)

Science Group	Capability Summary	$\lambda$	Spatial Res.	Spectral Res.	Field of View	Multiplex Gain	Primary Mode
Stars, Solar System, etc. →	Extrem AO -/ IFU & Pol.	0.9-2.5	SD > 0.9	30-300	3"	1	IFU + Pol + AO ←
Planets & Stars →	HIRES NIR Spectrometer → Optical HIRES Spec	0.9-2.5 0.3-1.0	See	70K (high stab) 50-100K		1	+ Molecular Absorber (X-disp)
<u>Star Formation / ISM</u>							
- How does Structure/Comp of ISM evolve?	①	8-17 $\mu$ m	DL	> 10 <sup>5</sup>	PS	1	X-dispersed MOS
- How are stars/planetary systems assembled?	②	1-5 $\mu$ m	DL	> 3x10 <sup>4</sup>	> 2'	100	X-dispersed
- What determines the masses of stars?	③	1-5 $\mu$ m	-	> 10 <sup>5</sup>	PS	1	contrast > 10 <sup>7</sup> ←
④	1-5 $\mu$ m	.05"		5	2"	1	
Structure/Evolution of MW + Nearby Galaxies ← cool millions	GLAO IR imager	3-100 $\mu$ m	nat. seeing	20K(?) 3K-40K(?)	> 10' 240'	~1000	imager. MOS ←
Galaxy Genesis	Optical MOS	9000 $\text{\AA}$					
Dark Matter Explorer	IR AO-fed spect.	2-3 $\mu$ m	0.05"	2K	20" x 20"	1	IFU
Stellar pops. Distant galaxies, high-z Universe	Optical IFU spec.	3-100 $\mu$ m	Nat.	5K	3' x 3'	1	IFU
DK energy + Gal. Form	Opt-NUV MOS	3400 $\text{\AA}$ - 1 $\mu$ m	(good) Natural seeing	> 500	> 30'	> 1000	MOS ←
1ST LIGHT + Gal. Form	(GLAO) fed NIR Mapper	NIR	0.2"	> 3000	10'	Quasi-Panoramic	IMAGER → dIFU → (TF?)





## \* Matter

- \* How do
- \* What is
- \* What is
- holes and

## \* Energy

- \* What is
- \* How did

## \* Life

- \* How con
- planets?
- \* How do
- \* How do
- blocks of



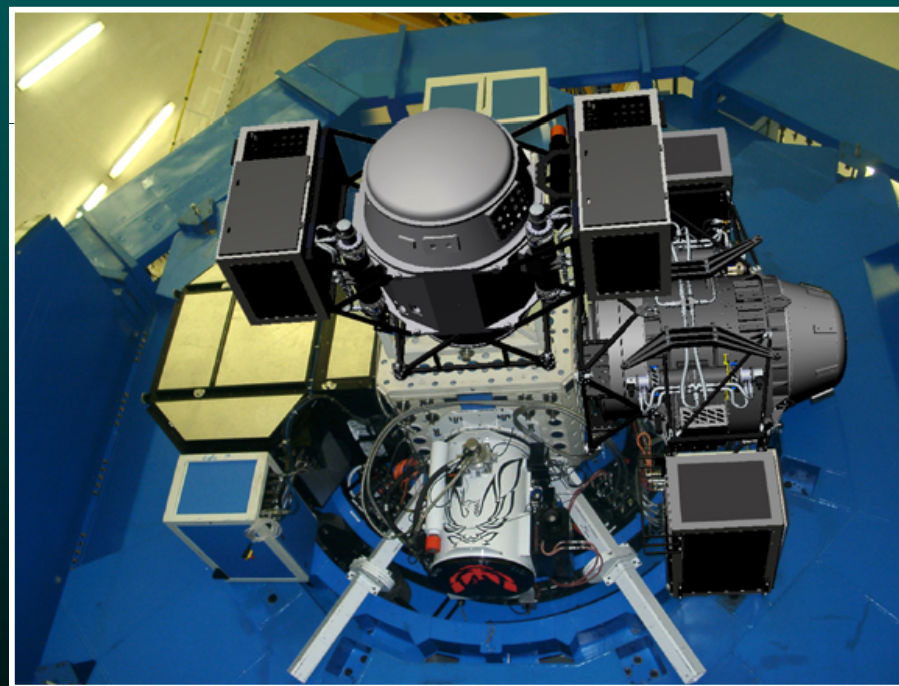
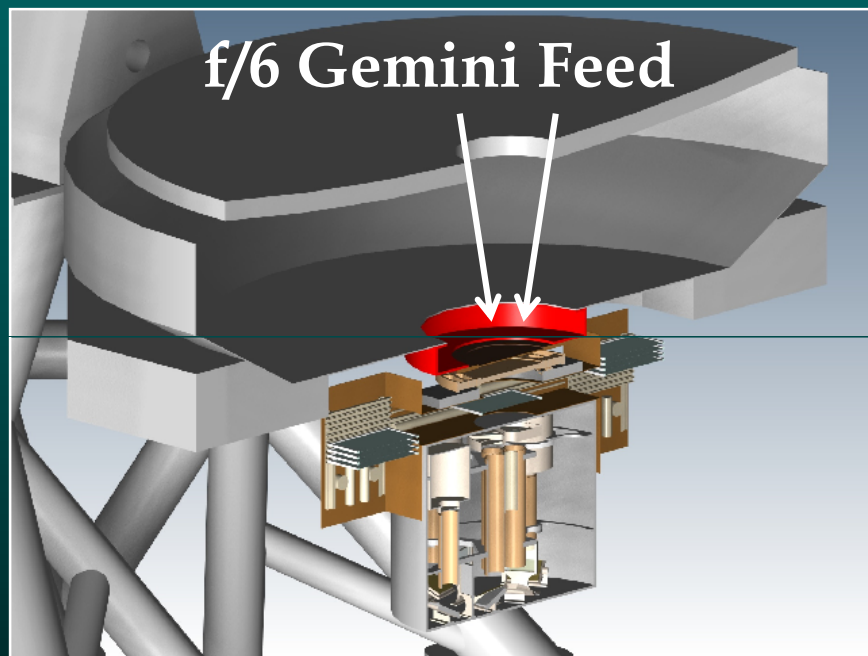
... scales?

...ive black

...ng earth-like

...nical building

Wide Field Optical MOS Concept



HRNIRS Concept on up-looking port



<b>New Instruments</b>	<b>Total Estimated Cost (USD)</b>
Extreme AO Coronagraph	12,961,313
High Resolution NIR Spectrometer	17,260,000
High Resolution Optical Spectrometer	8,495,000
High Resolution MIR Spectrometer	20,481,250
MCAO fed NIR MOS	24,303,750
Wide Field Optical MOS	31,159,375
AO-fed NIR Spectrometer	14,876,625
IFU Optical Spectrometer	14,000,000
GLAO NIR Imager	14,139,375
GLAO NIR Spectrometer	21,592,500
Wide Field Fiber-Fed Optical MOS	31,977,125
<b>New Facilities</b>	
f/6 Telescope Configuration (1 telescope)	33,035,625
GLAO System including Adaptive Secondary	23,535,960



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## *The WFMOS Breakthrough...*

---



**An important meal at Hale Pohaku**



# *Assembling the Many Pieces of the WFMOS "Puzzle"*





# *Assembling the Many Pieces of the WFMOS "Puzzle"*

Chris Packham visited Tohoku, NAOJ, Tokyo, ISAS/JAXA, Hiroshima, Kobe, Kyoto to discuss Subaru/Gemini collaboration...



Attended 3 Annual Subaru User's Meetings

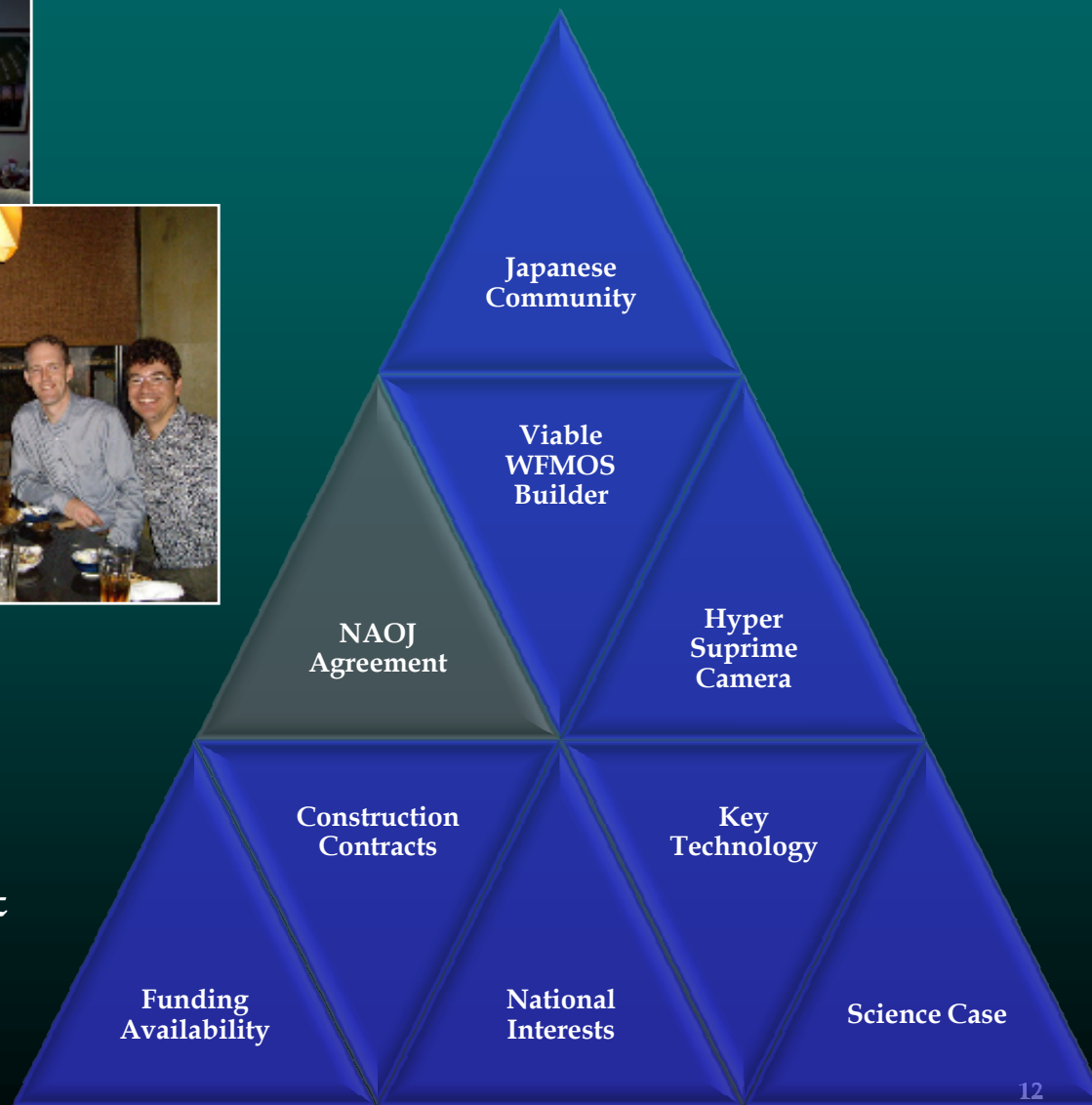




# Assembling the Many Pieces of the WFMOS "Puzzle"



Over a dozen meetings in Hilo & Tokyo to negotiate the shared development/use of WFMOS...



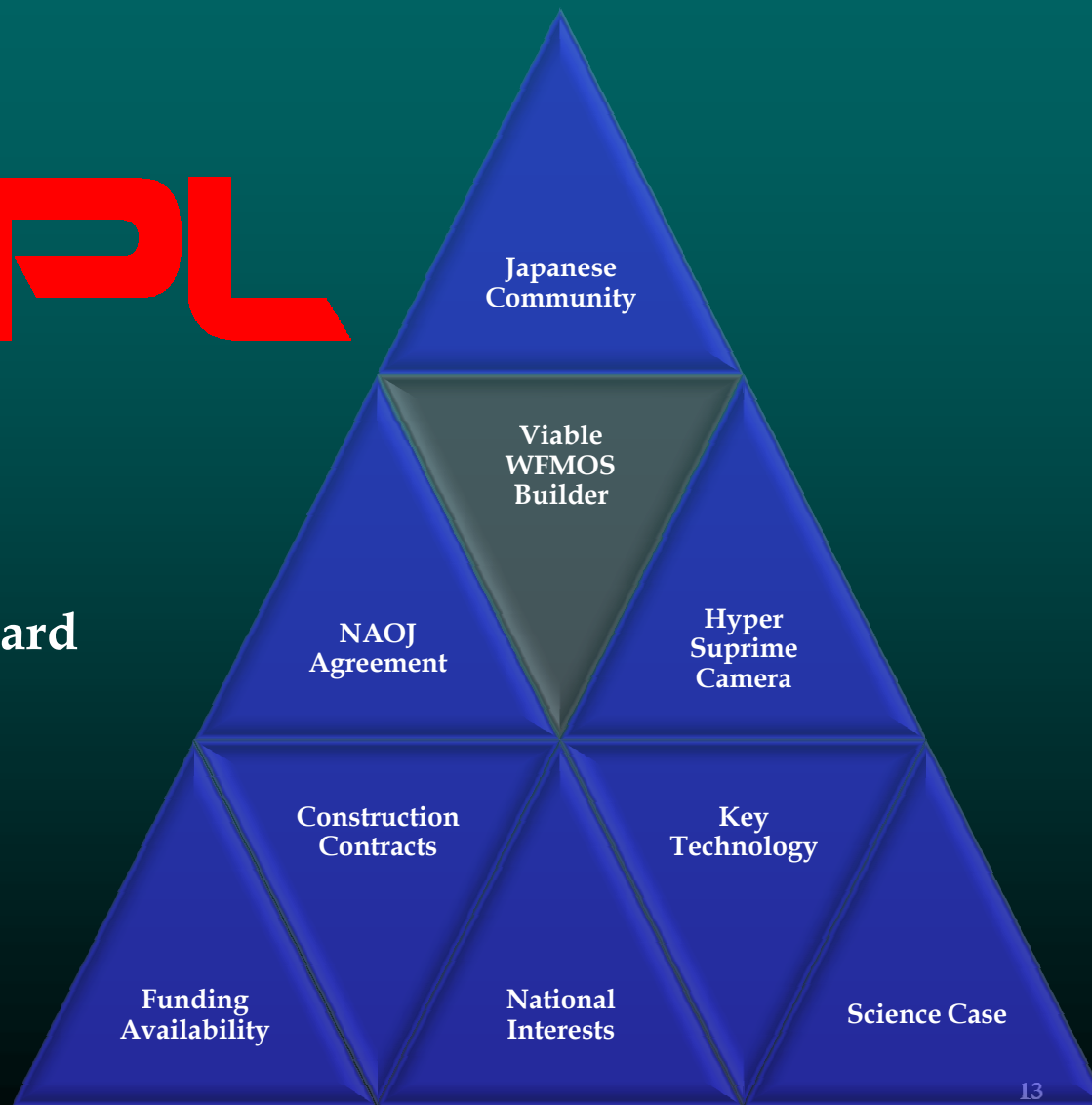


# *Assembling the Many Pieces of the WFMOS "Puzzle"*



**JPL & AAO Compete**

**Recommend Going Forward with JPL Led Team**

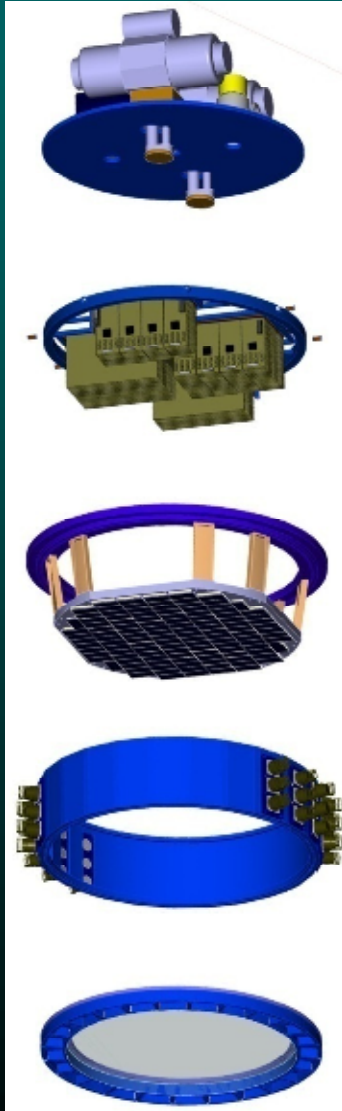




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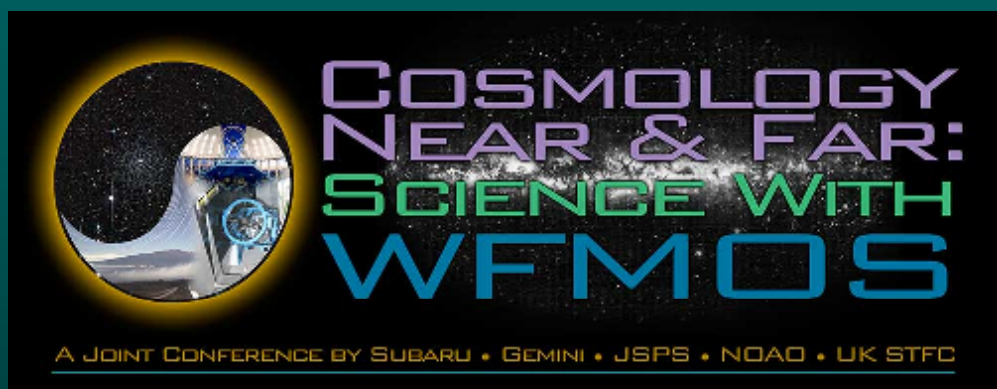
# *Assembling the Many Pieces of the WFMOS "Puzzle"*

## Hyper Suprime Camera Assembly

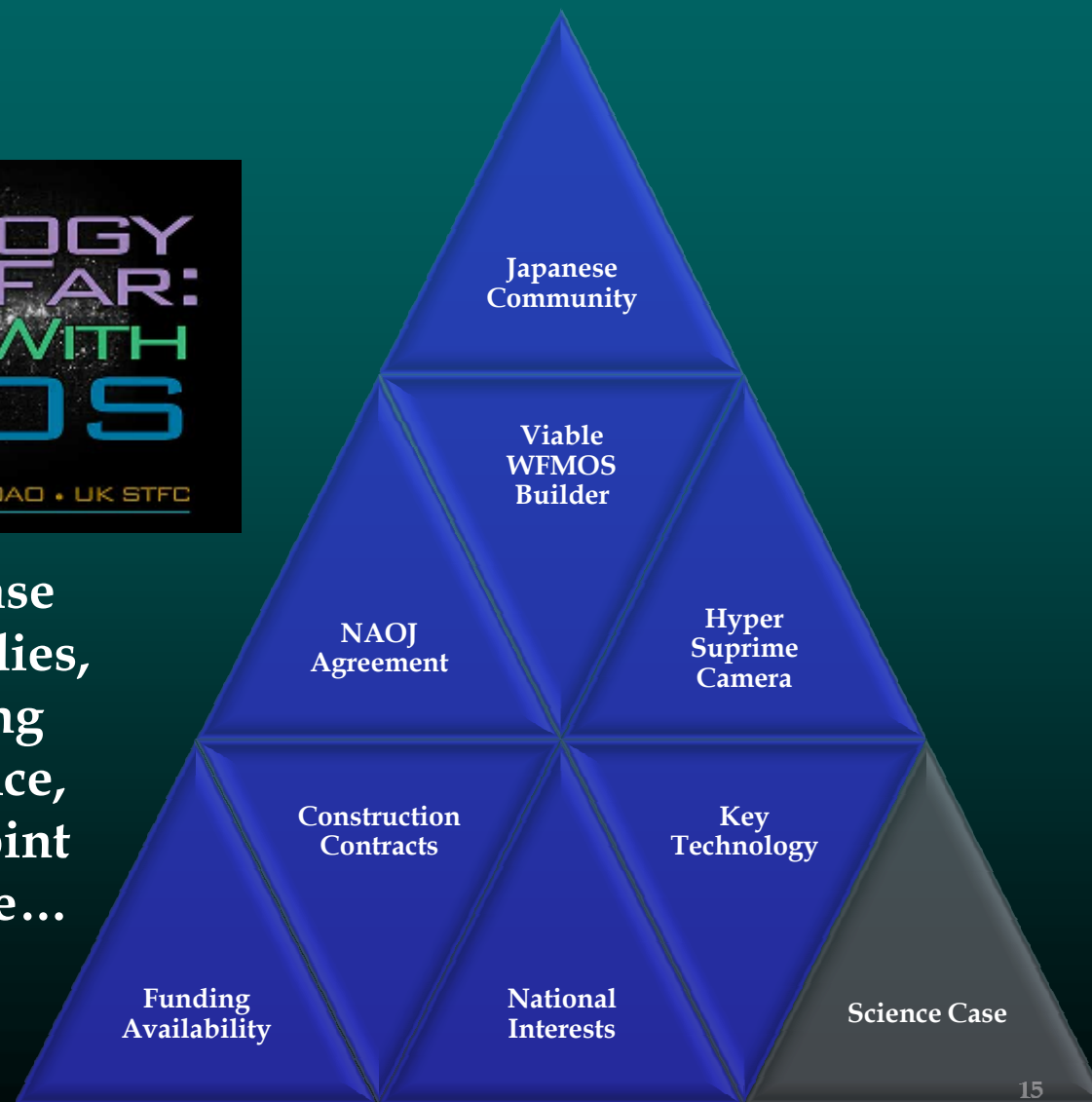




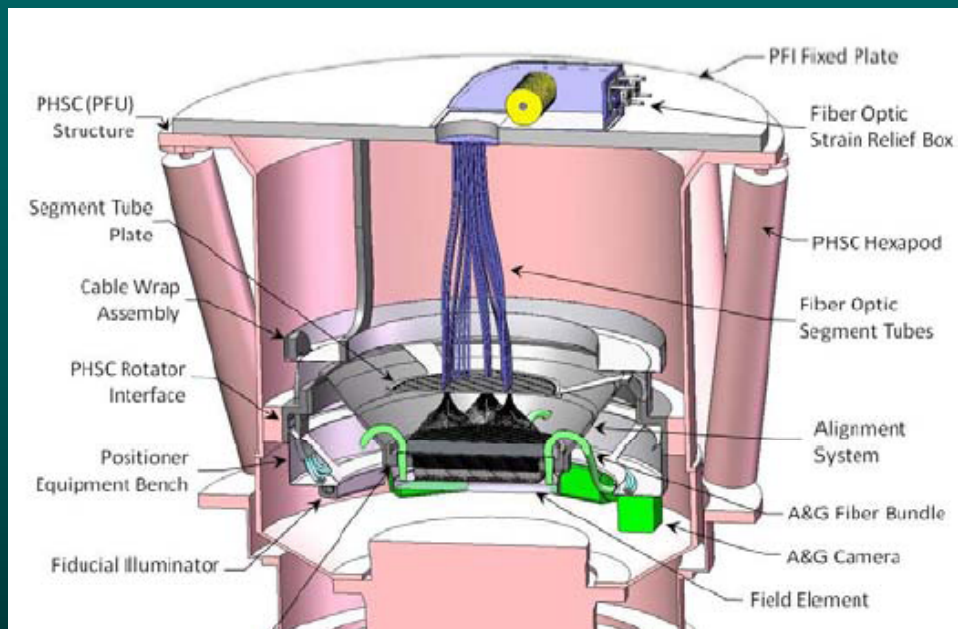
# *Assembling the Many Pieces of the WFMOS "Puzzle"*



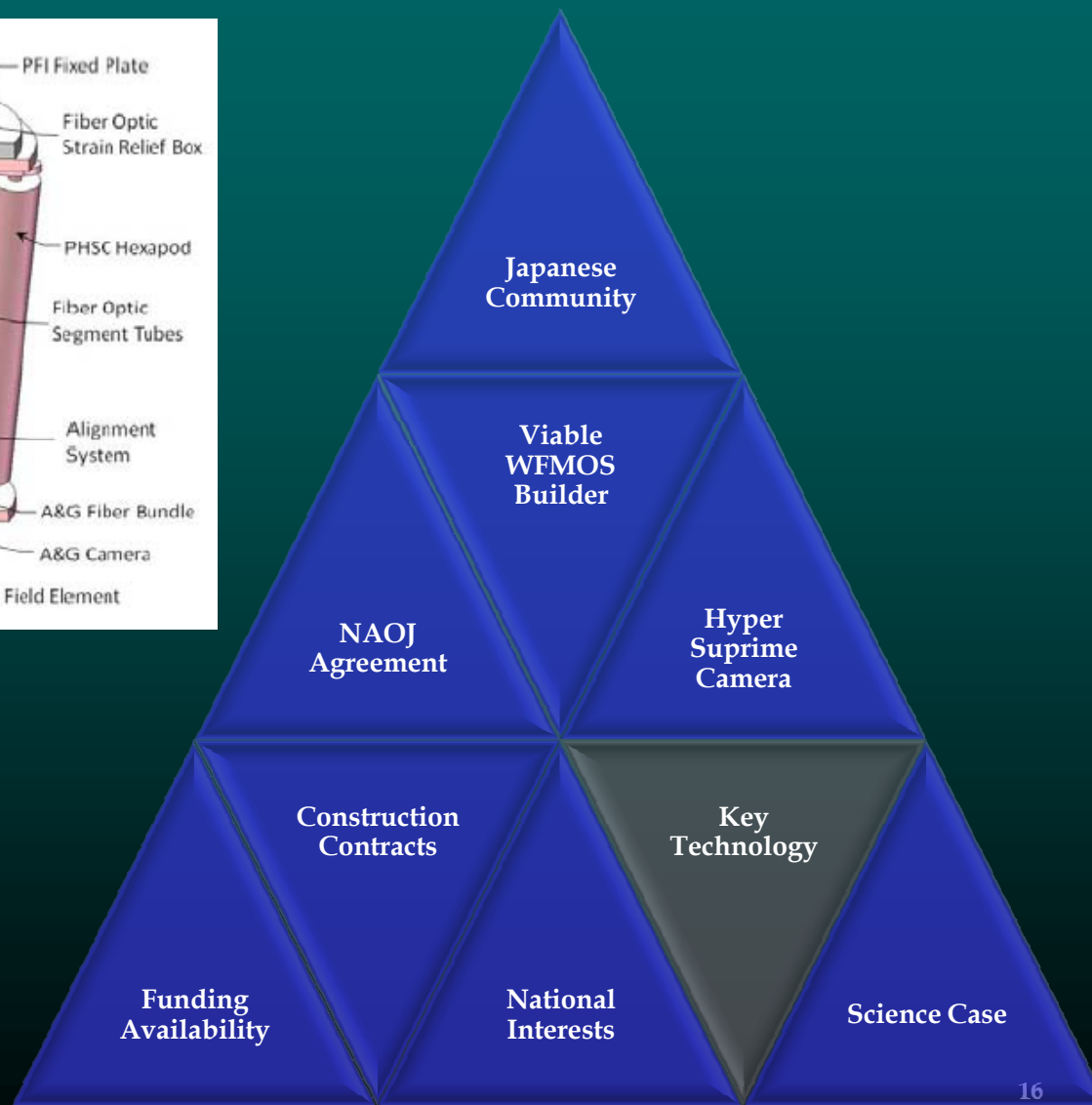
Robust WFMOS science case developed via 2 funded studies, 2009 Subaru User's Meeting dedicated to WFMOS science, and 2008 Gemini/Subaru joint WFMOS science conference...



# *Assembling the Many Pieces of the WFMOS "Puzzle"*



**Fiber Positioner**





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# Assembling the Many Pieces of the WFMOS "Puzzle"

Final Report of the Committee  
on  
Access to Large Telescopes for Astronomical Instruction and Research  
(ALTAIR)<sup>1</sup>

Submitted March 9, 2009

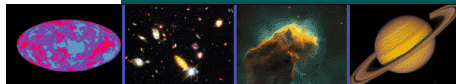
**Committee Members**

- Daniel Eisenstein, University of Arizona
- Heidi Hammel, Space Science Institute
- Lynne Hillenbrand, California Institute of Technology
- Christopher Johns-Krull, Rice University
- David Koo, University of California at Santa Cruz
- Nancy Levenson, University of Kentucky
- Tom Matheson, National Optical Astronomy Observatory
- Andrew McWilliam, Carnegie Observatories
- John Munnier, University of Michigan
- Joan Najita, (vice-chair) National Optical Astronomy Observatory
- Lisa Prato, Lowell Observatory
- Larry Ramsey, (chair) Penn State University
- Seth Redfield, University of Texas at Austin, Wesleyan University
- Tom Soifer, Spitzer Science Center

**Liaisons and Observers**

- Tom Barnes, NSF observer
- Michael Briley, NSF observer
- Eileen O'Connell, NSF observer

## Astro 2010



CANADIAN ASTRONOMY AND ASTROPHYSICS IN THE 21ST CENTURY



### THE ORIGINS OF STRUCTURE IN THE UNIVERSE

Report of the NRC-NAS/NSF Long Range Planning Panel



## Canada



## UK/Europe



## Australia





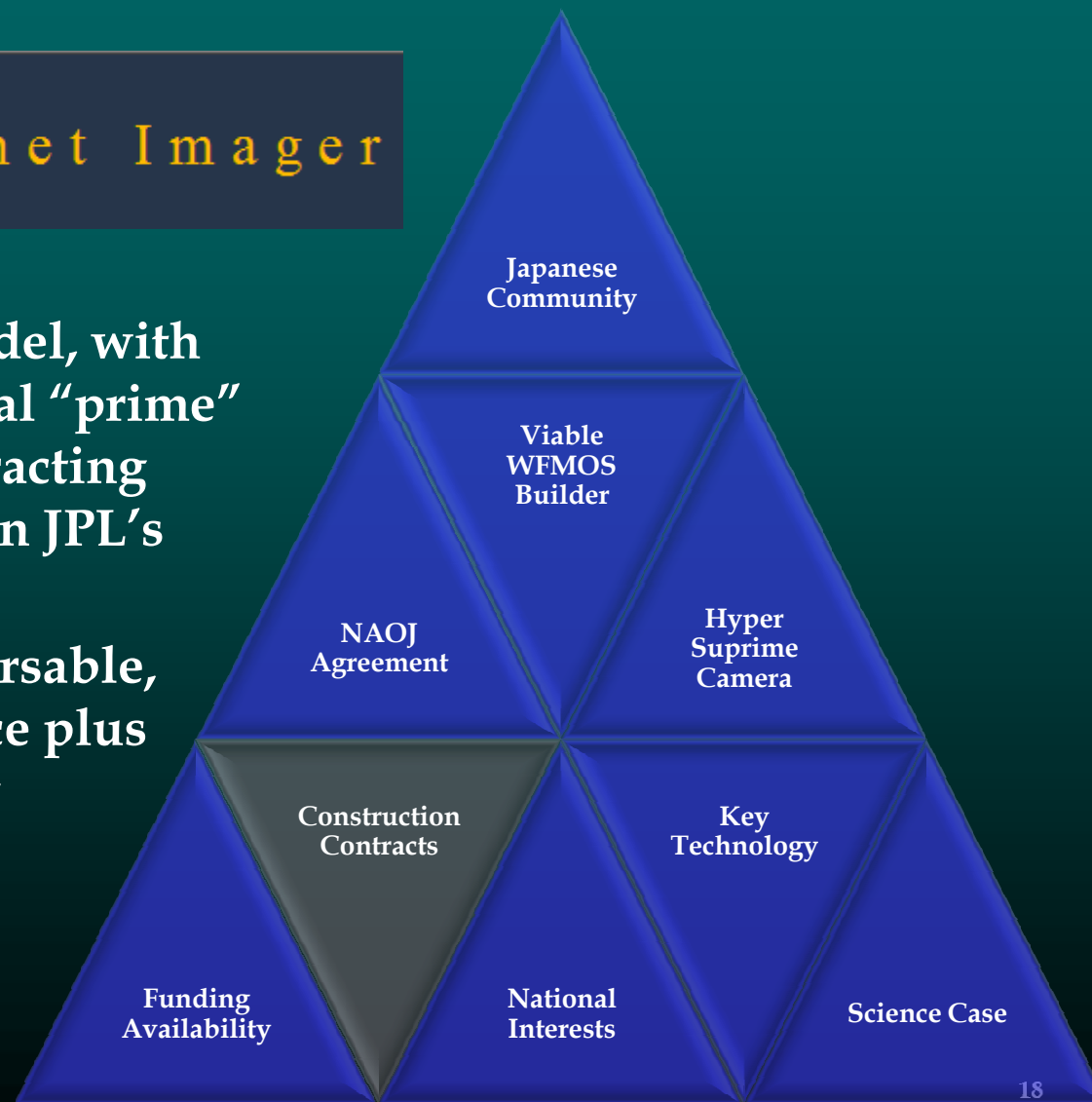


# *Assembling the Many Pieces of the WFMOS "Puzzle"*



Gemini Planet Imager

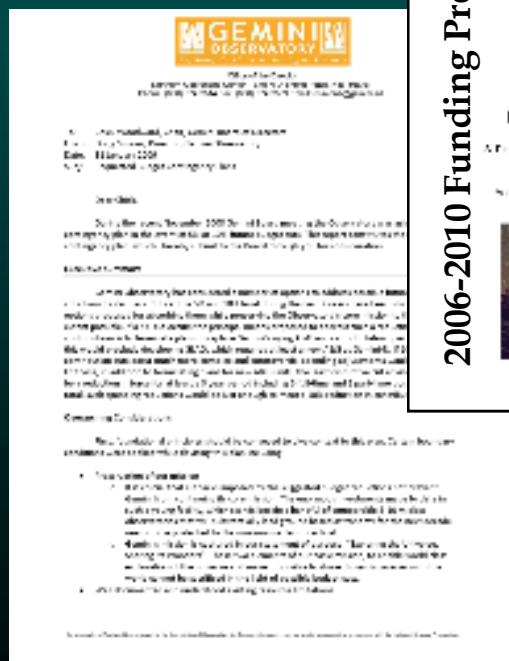
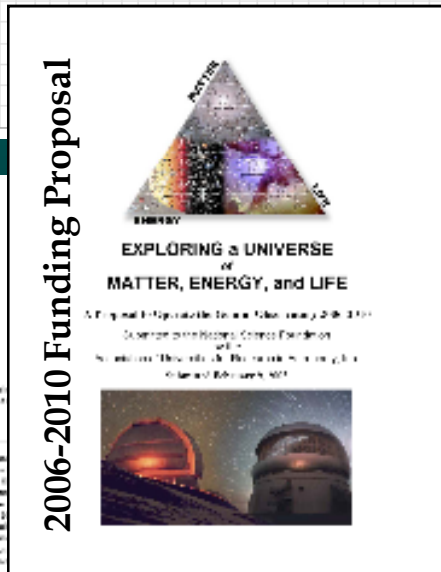
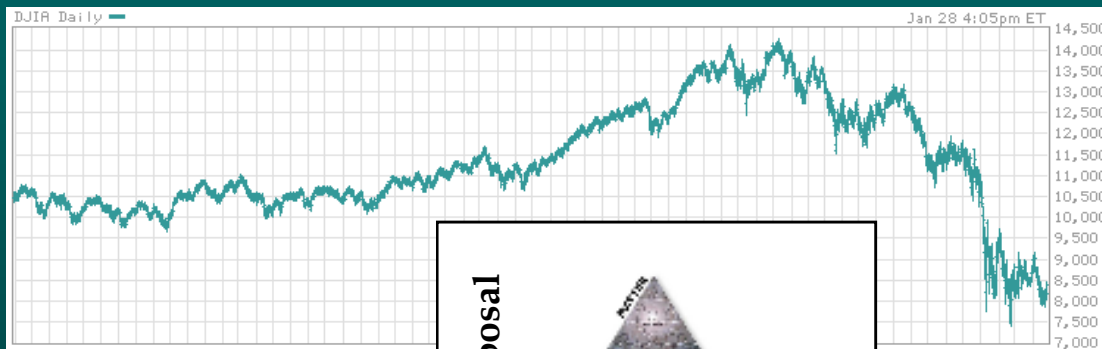
- \* JPL proposing similar arrangement to GPI model, with JPL acting as a functional "prime" and Gemini direct contracting with other institutions in JPL's team
- \* JPL portion cost reimbursable, remainder are fixed price plus shared risk contingency





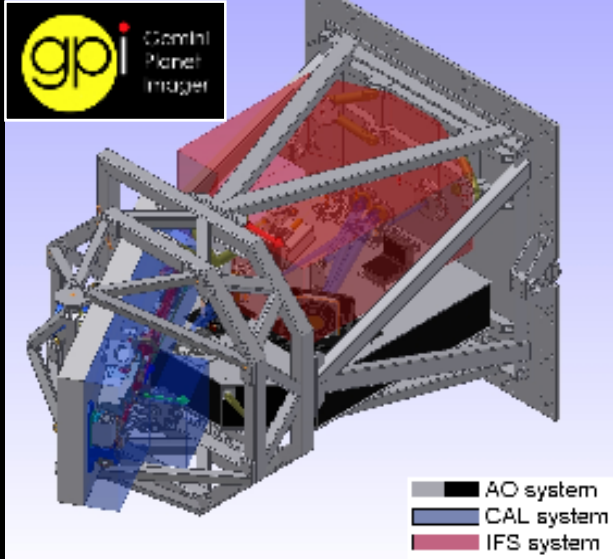
# Assembling the Many Pieces of the WFMOS "Puzzle"

## US Stock Market Performance Over Last 5 Years



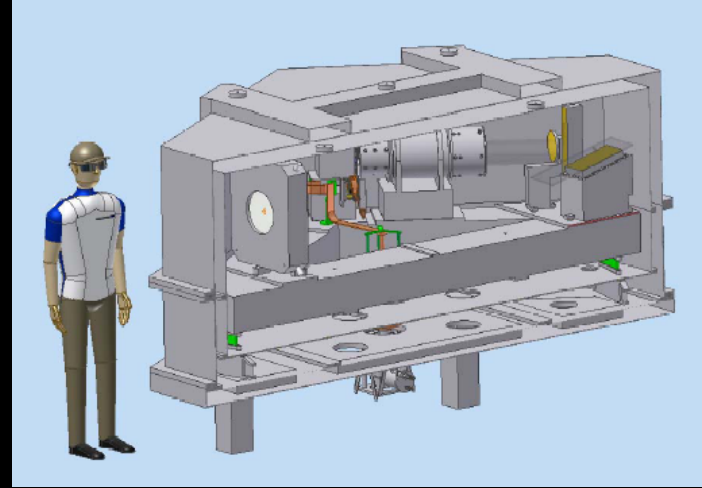
Gemini Contingency Plan



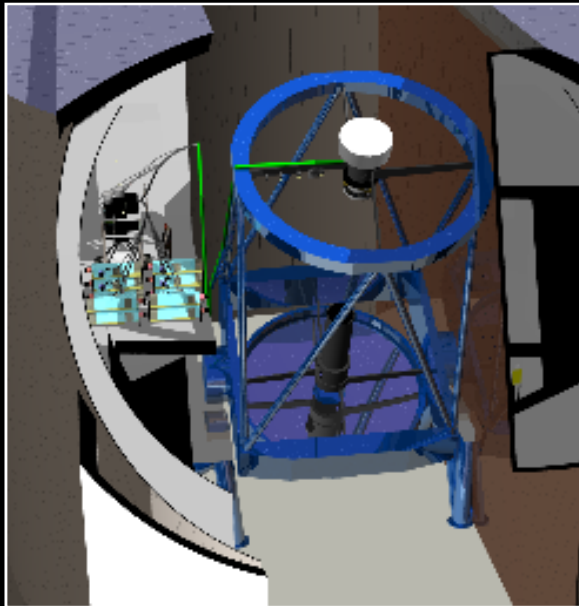


**Gemini  
Planet  
Imager**

**Precision  
Radial  
Velocity  
Spectrometer**

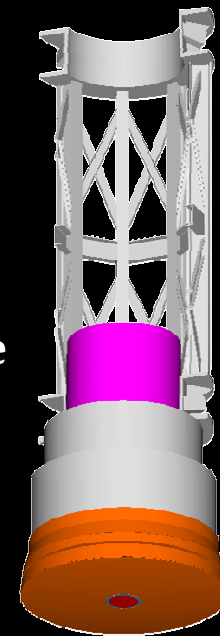


# *Aspen Instrument Program*



**Wide  
Field  
Multi-  
Object  
Spectrometer**

**Ground  
Layer  
Adaptive  
Optics**





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*From Gemini's Board of Directors  
Meeting (May 12-13)...*

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**“The Board acknowledges receipt of the results of the WFMOS design studies in a report from the Observatory. The Board also acknowledges receipt of a draft of the Memorandum of Understanding between NAOJ and the Gemini Partnership describing the obligations and expectations of implementing and operating the WFMOS instrument on Subaru, the principles guiding the execution of WFMOS science (both campaign and PI driven) opportunities, and the mechanism for time exchange and access to the Gemini telescopes both North and South.”**



*From Gemini's Board of Directors Meeting (May 12-13)...*

---

**“The Board recognizes there is strong science support in the Gemini community for the development of WFMOS, intense interest in developing an effective collaboration with NAOJ, and strong science motivation for the project. However, WFMOS has significantly increased in cost, and the Gemini Partnership can identify neither the necessary resources to proceed to WFMOS PDR, nor the totality of funds required in the out years to successfully complete the instrument. The Board regretfully terminates the WFMOS Aspen initiative.**

**The Board unequivocally endorses the GPI project as the completion of the Aspen program.”**



*From Gemini's Board of Directors Meeting (May 12-13)...*

---

**“The Board wishes to express its deep appreciation for the extraordinary effort of the Gemini Observatory, our partner communities, as well as our colleagues at NAOJ. The Board does recognize the engagement between the Gemini partnership and NAOJ and the communities they represent as being enormously beneficial to the expanding international nature of astronomy, and more importantly, the development of strong friendship and aspiration for collaborations to advance science. The Board greatly values the developing scientific partnership between the Gemini and Subaru communities, as manifest in the first ever joint Gemini/Subaru science conference in May 2009, and remains open to exploring areas of mutual interest and benefit to further strengthen this partnership.”**

- \* Aspen was a “blue sky” approach to developing a strategic plan
- \* Intent of Aspen was to -
  - \* Capitalize quickly on the investments made in new 8 m telescopes
  - \* Establish scientific legacy early for these facilities, before the era of JWST and ELT’s
- \* Aspen produced instrument concepts that were vastly more expensive than the first generation of 8 – 10 m class instruments
  - \* A “moon shot” by previous instrument standards
- \* A \$75M “planning budget” was established and “pay as you go” approached adopted rather than firm funding commitments
- \* In practice, funding provided ~1/2 as fast as needed, which slowed down the program badly, yet astronomy moved on...
- \* Global economic crisis, combined by a Gemini Partnership that is in a state of transformation (International Agreement expiring), arrived at the worst possible time



Collaboration





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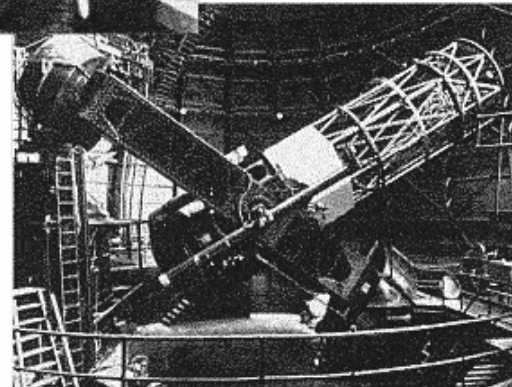
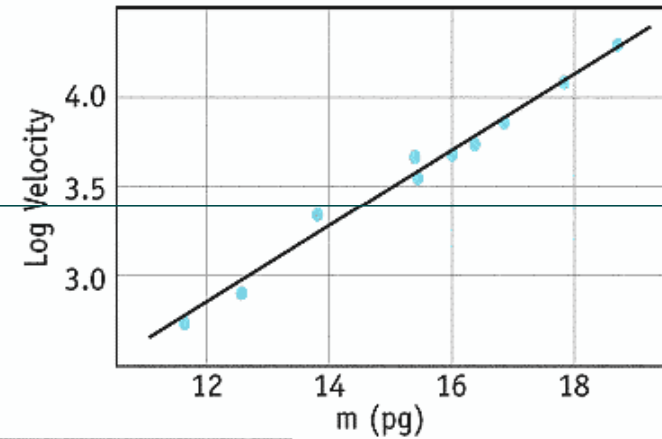
# *Gone are the Days...*



## DISCOVERY OF EXPANDING UNIVERSE



Edwin Hubble



Mt. Wilson  
100 Inch  
Telescope



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# *The Globalization of Astronomy is Upon Us*

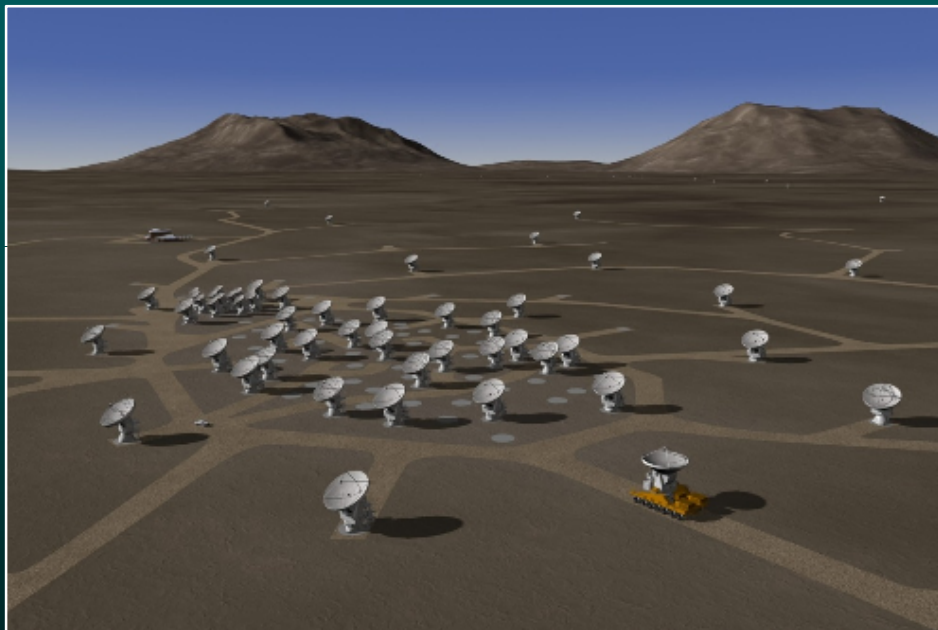


\*Table prepared by Dennis Crabtree, Gemini Observatory

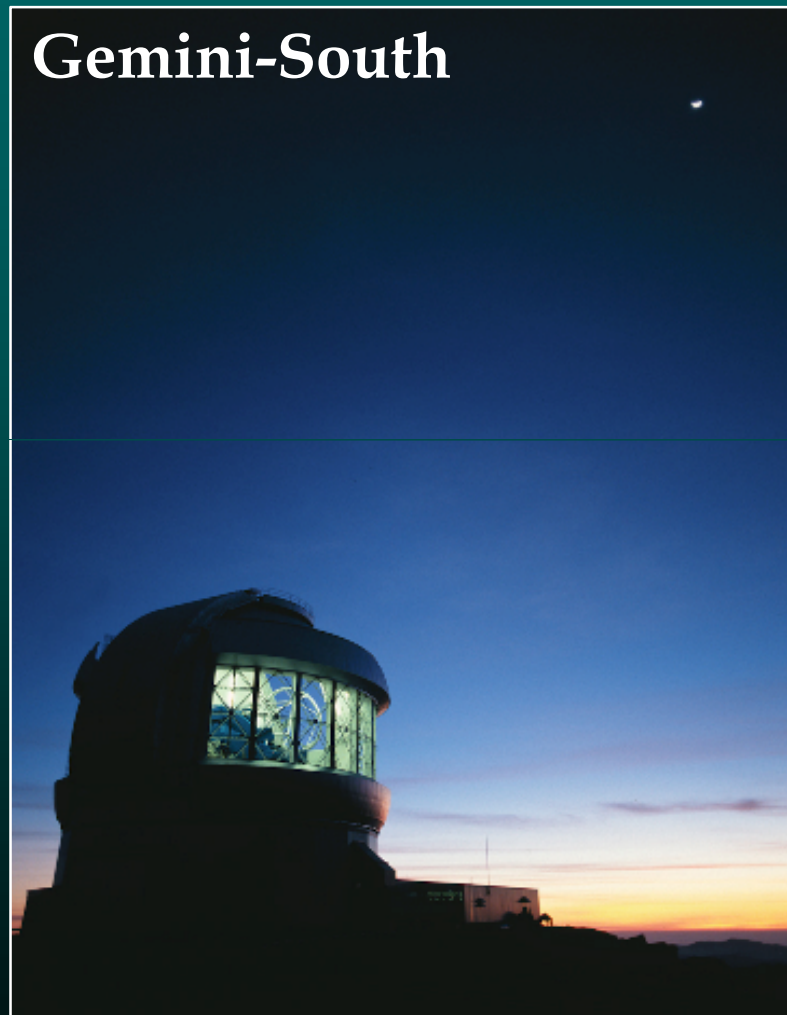


# *Synergies Will Drive Us Together*

## ALMA



## Gemini-South





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# *Emergence of a Pan-Pacific Astronomy Alliance*





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## *Where From Here?*

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- \* In ~4 months, Gemini will submit its next 5 yr budget proposal to its funding agencies
- \* A key part of this proposal will be a requirement that Gemini's future instrumentation funding commitment must be at the same level as our base operations funding
- \* In this proposal, Gemini will focus on next-gen instruments designed to –
  - \* Replace an aging fleet of workhorse instruments
  - \* Maintain a balance between a blend of workhorse instruments and a few advanced “pathfinder” instruments designed to take advantage of Gemini's design strengths

*Gemini will continue to work closely with Subaru to*



*recognizing that even Subaru Telescope  
WFEMOS, we have*

*simply come too far down the road together to turn  
back now...*

## WFEMOS as a Pathfinder into the Future of Astronomy